

*Accelerating the simulations of nonlinear dynamical systems with deep learning*

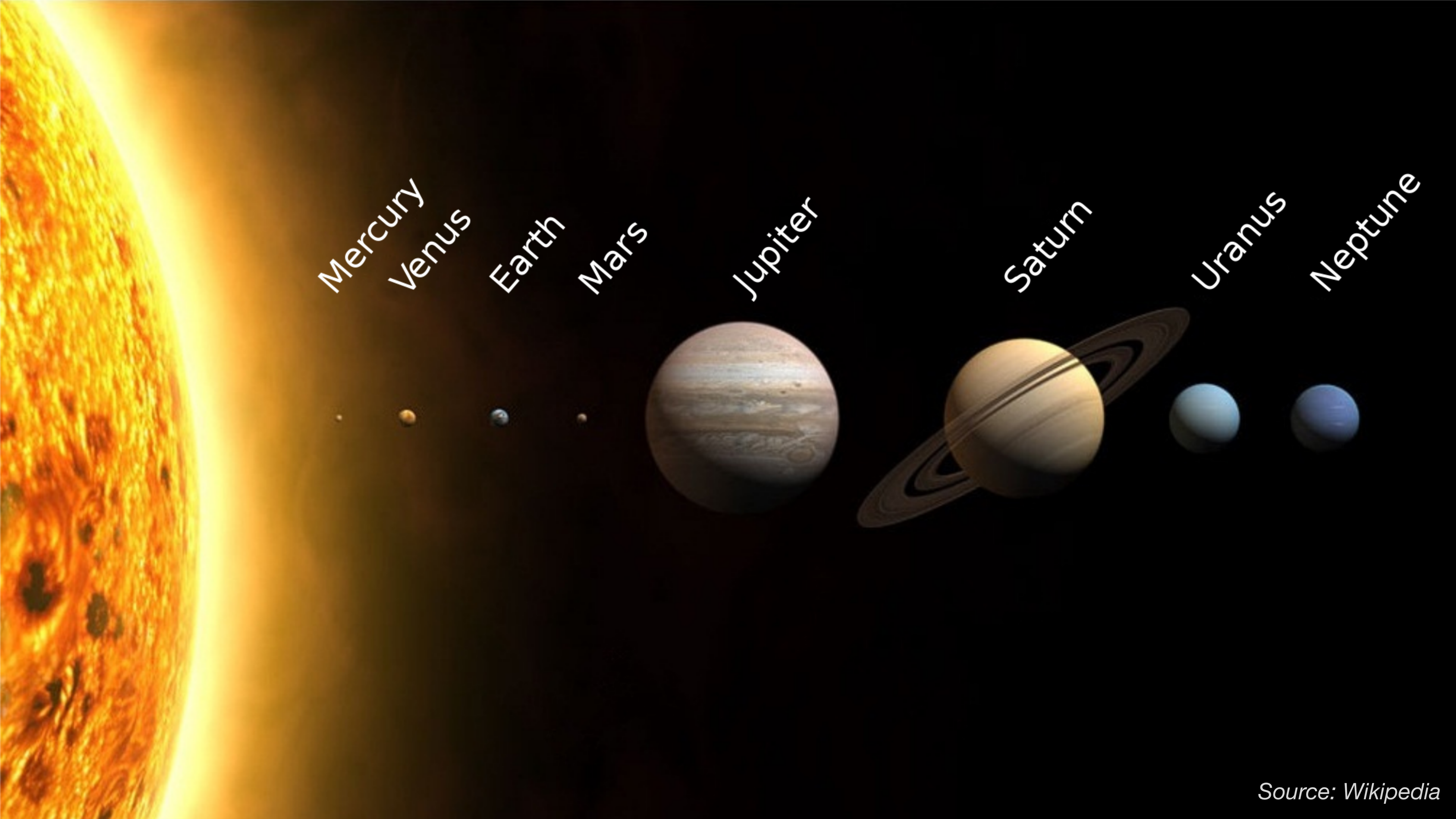
Maxwell Cai (Leiden U/SURF)

Simon Portages Zwart (Leiden U)

Damian Podareanu (SURF)

Valeriu Codreanu (SURF)

Caspar van Leewuen (SURF)



Mercury  
Venus

Earth

Mars

Jupiter

Saturn

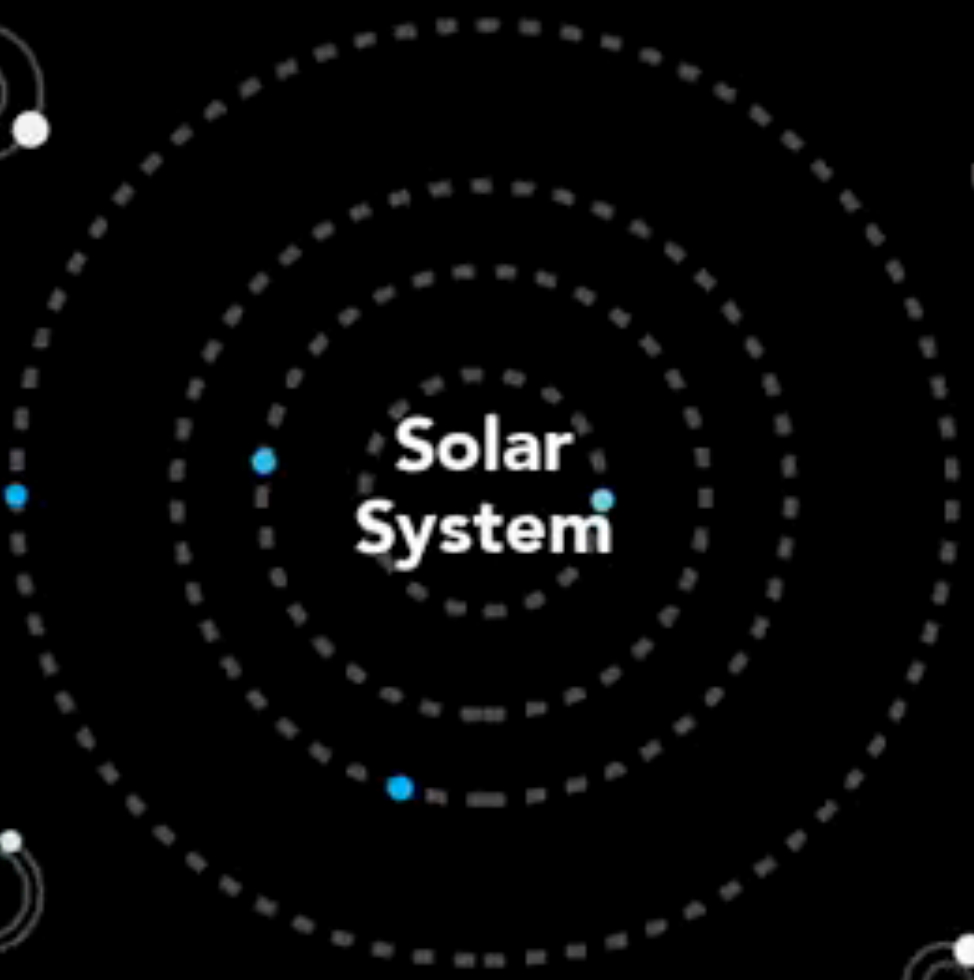
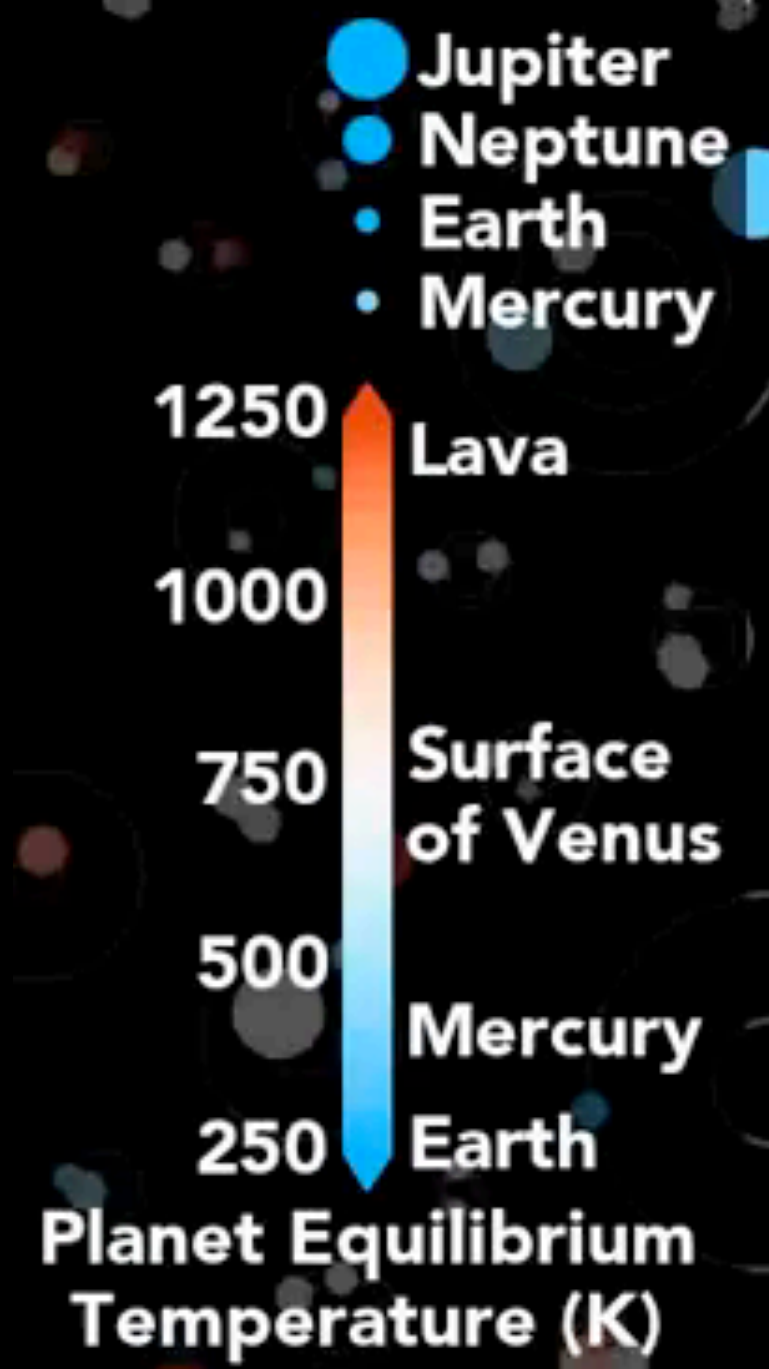
Uranus

Neptune

*Credit: NASA*



Kepler Orrery IV  
23 Nov 2010  
By Ethan Kruse  
@ethan\_kruse

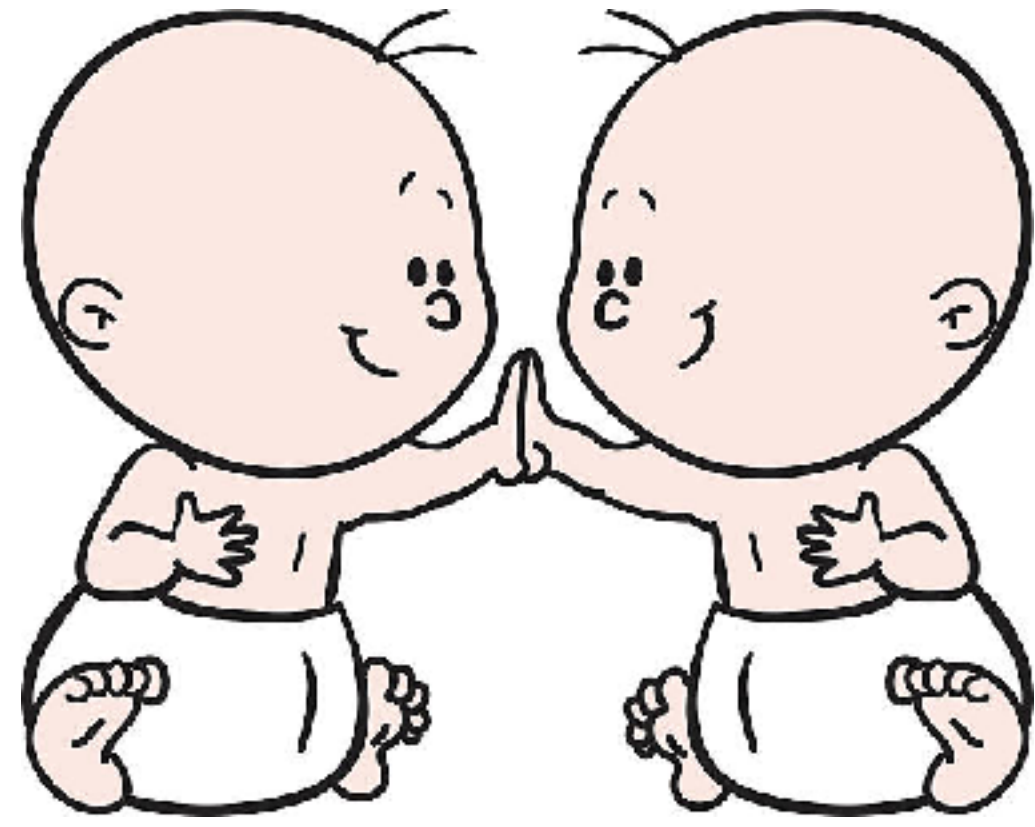


4,082 planets   3,046 planetary systems   660 multiple planetary systems   (15 June 2019)

*Credit: NASA*

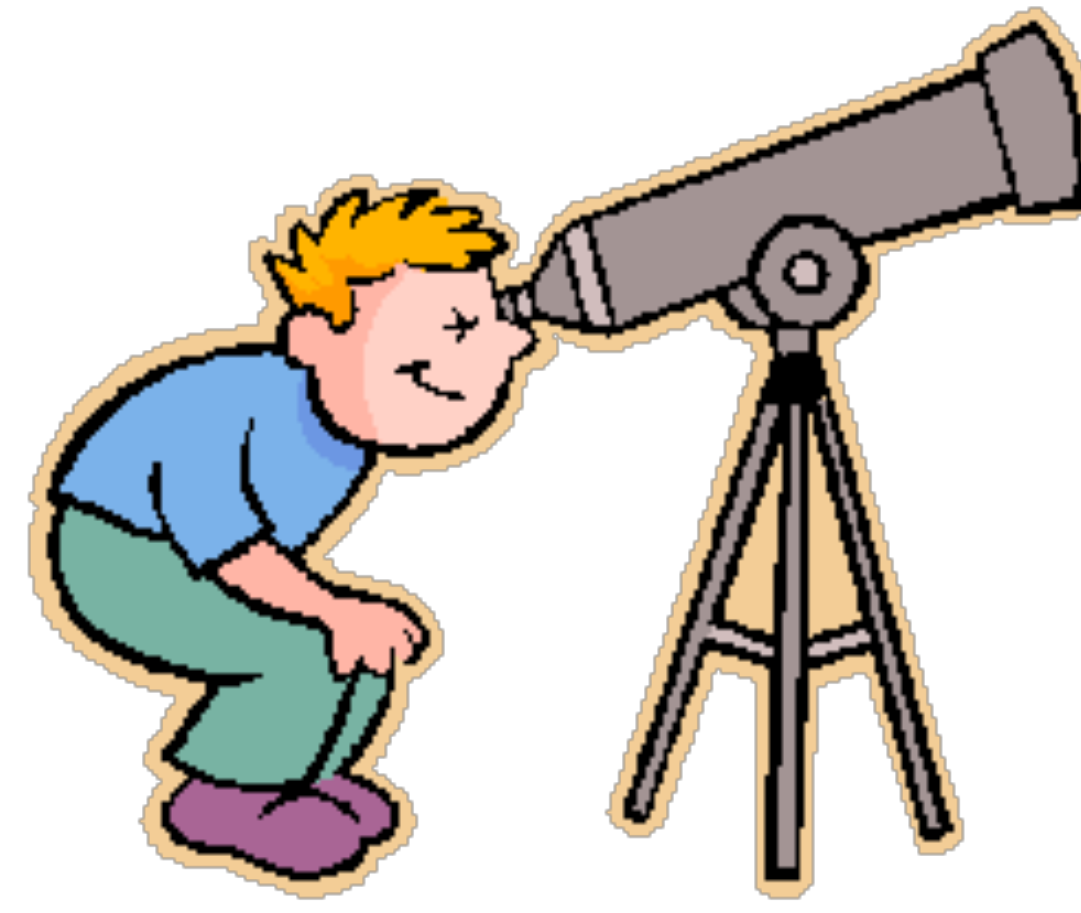


# METHODOLOGY



Identical twins

Different education  
Different environments



Astronomer

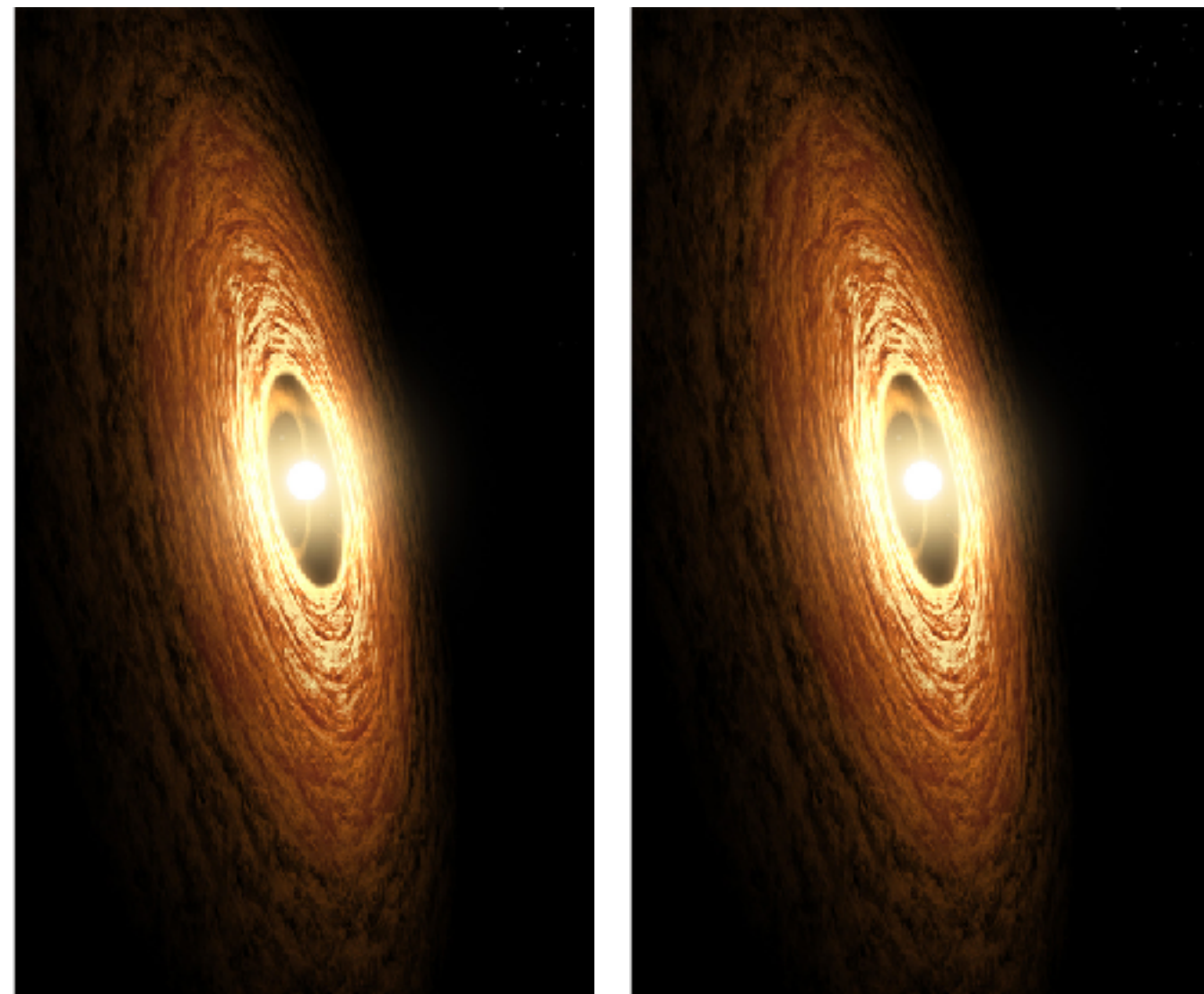


Musician

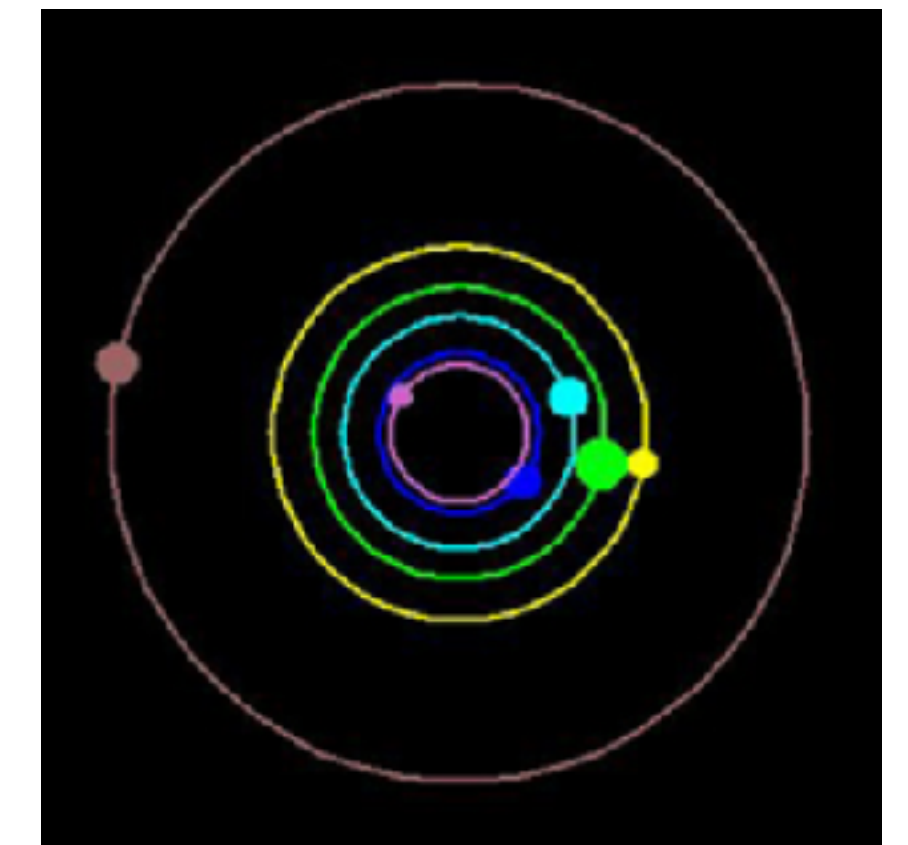
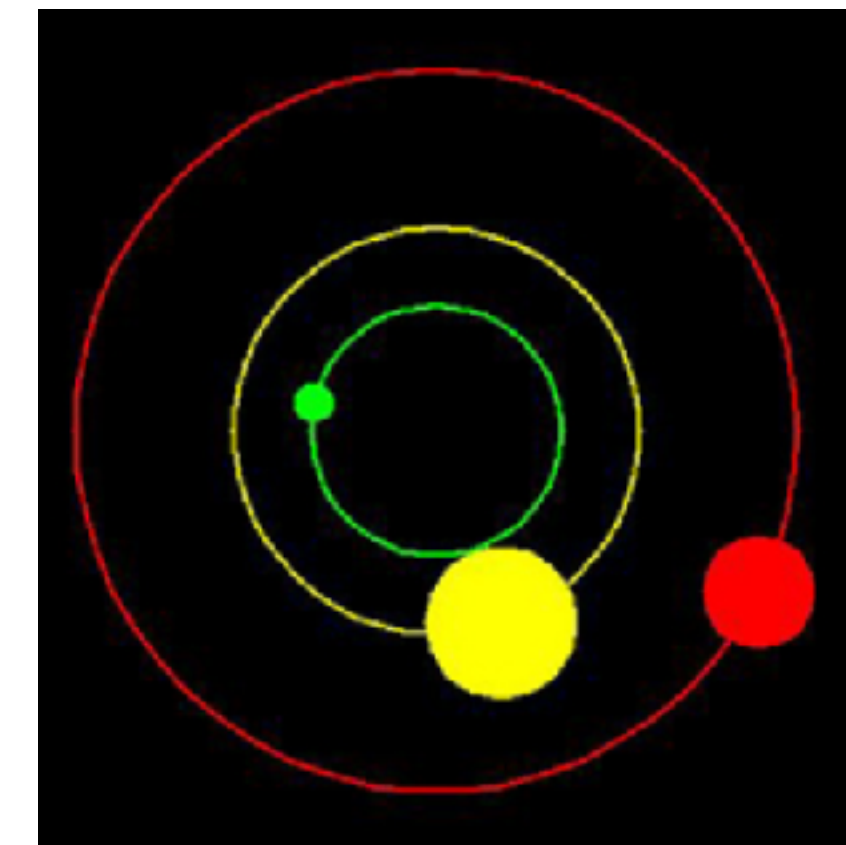
*Less Diverse*

TIME

*More Diverse*



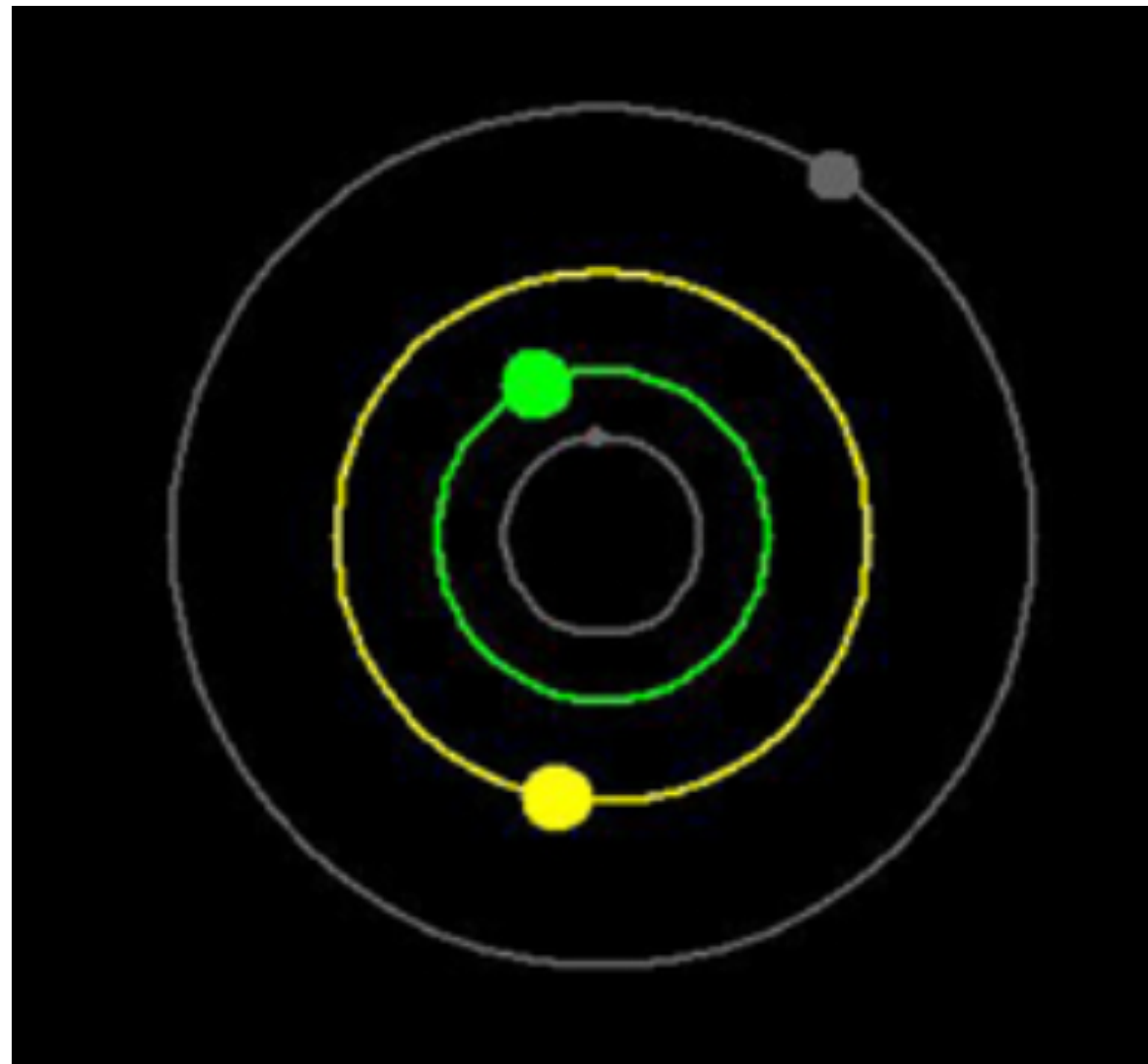
Different environments



Different orbital architectures

# Multi-scale Modeling

**Solar System**



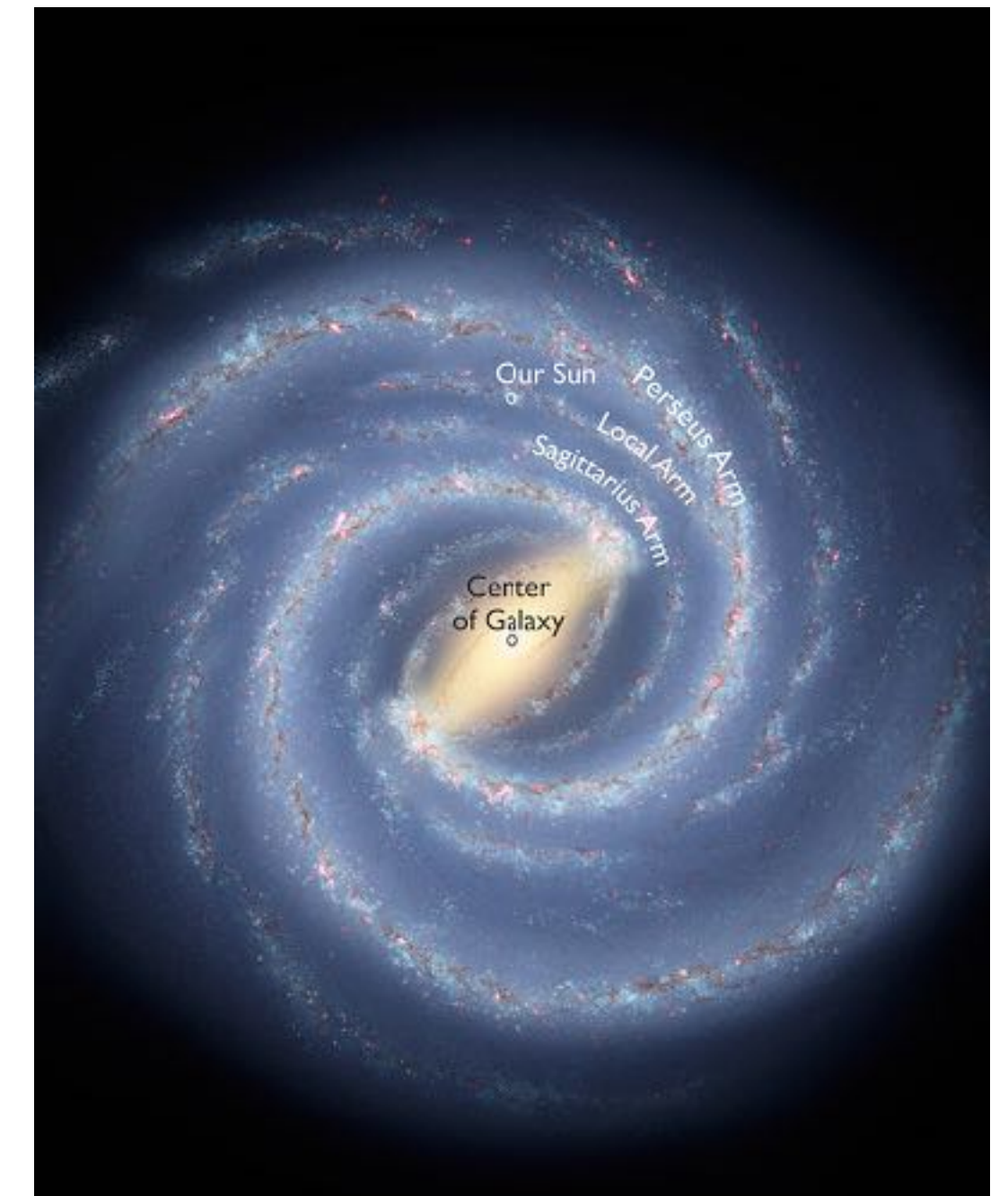
**1-2 stars  
A few planets  
100 AU**

**Star cluster**



**1000 - 1000000 stars  
100,000,000 AU**

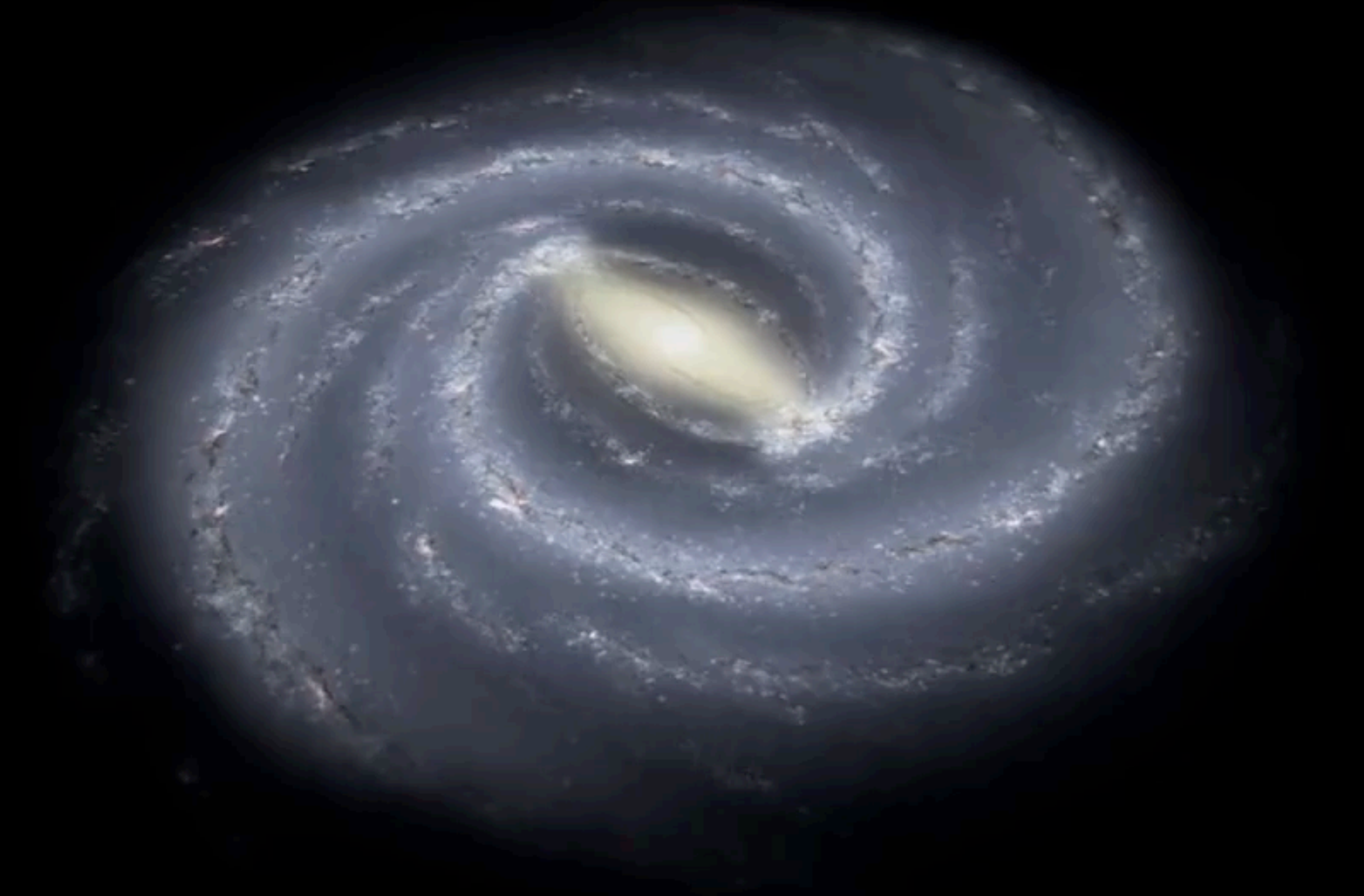
**Milky Way**



**200,000,000,000 stars  
100,000,000,000,000 AU**

**1 AU = distance from Sun to Earth = 150,000,000 km**

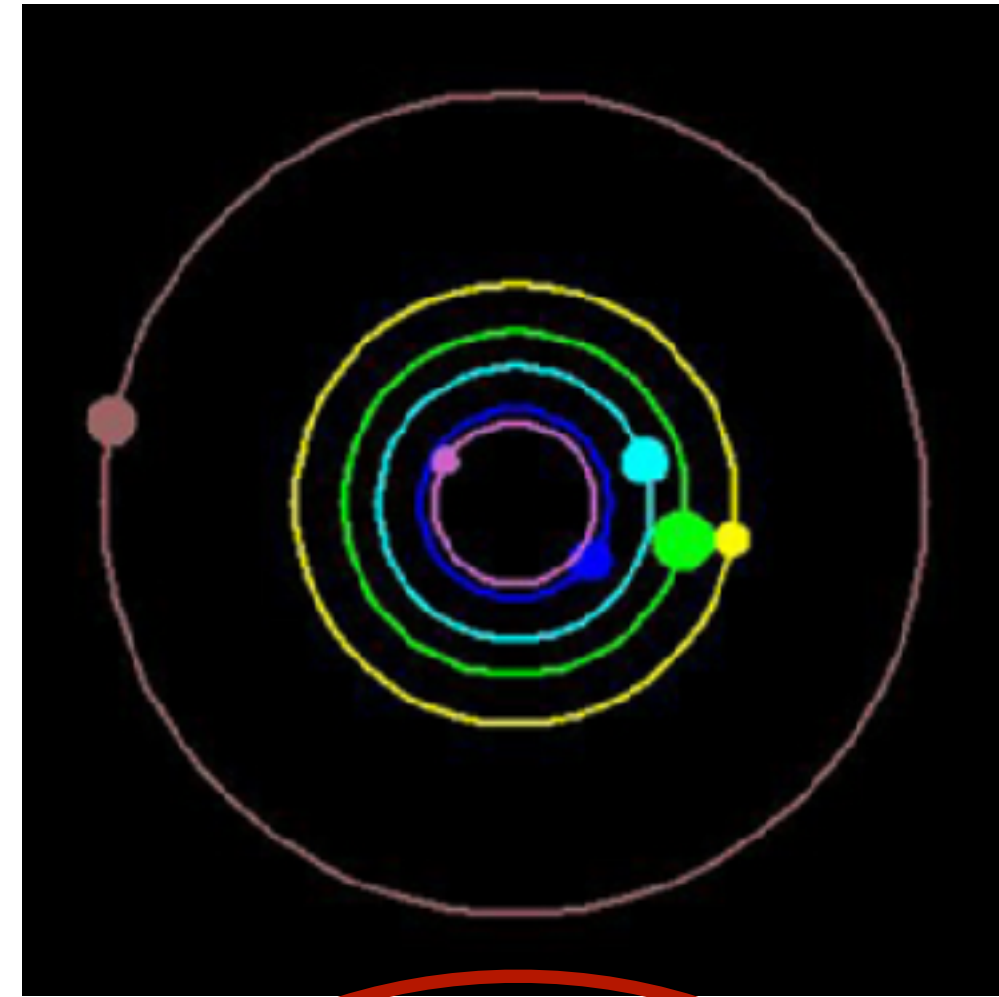
149 975 Light Years



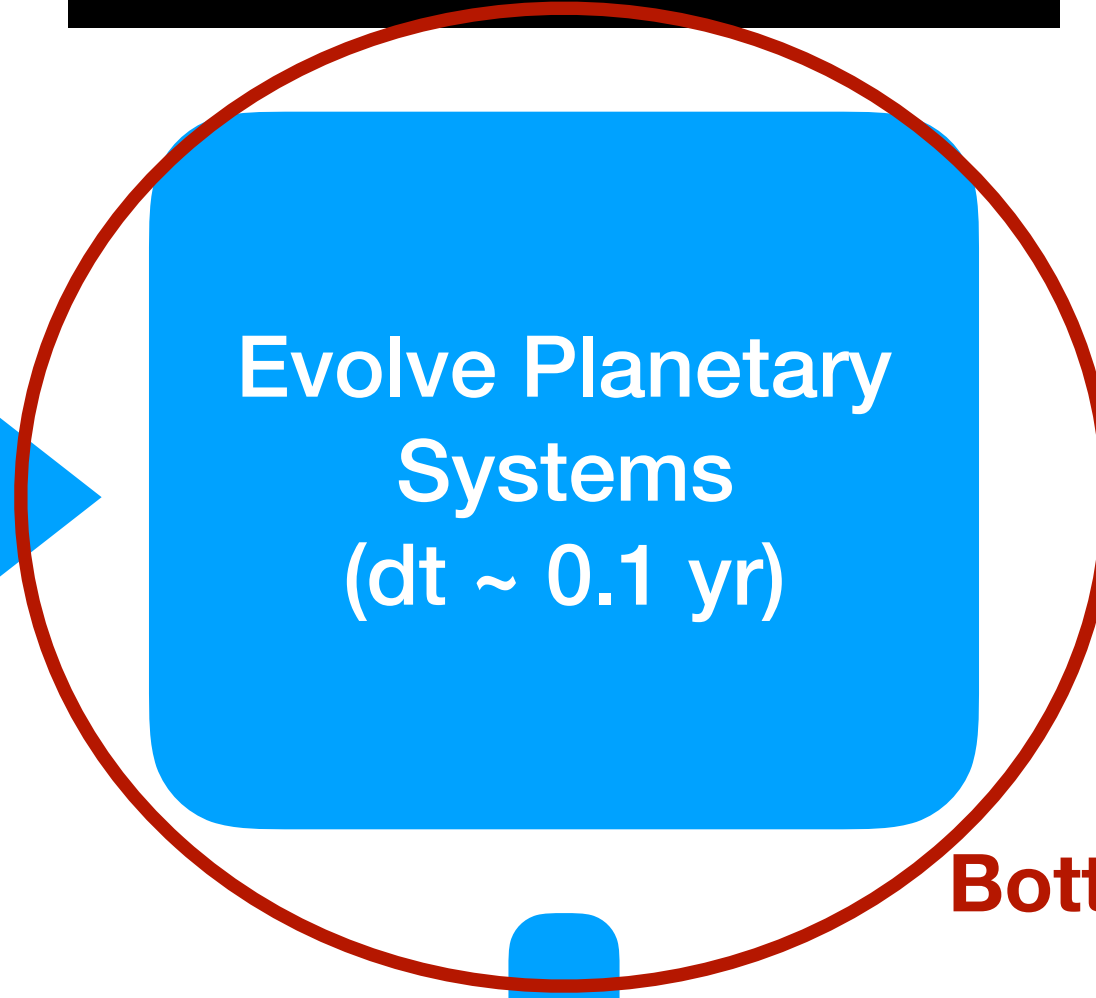
The Milky Way,  
the galaxy in which our sun goes round every 200 million years.



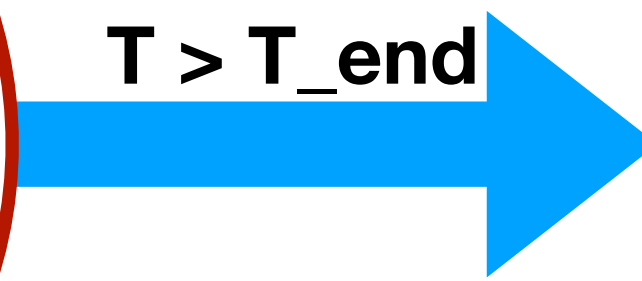
# Multi-scale Modeling



Evolve Star clusters  
(dt ~ 1000 yr)



Evolve Planetary  
Systems  
(dt ~ 0.1 yr)



Done

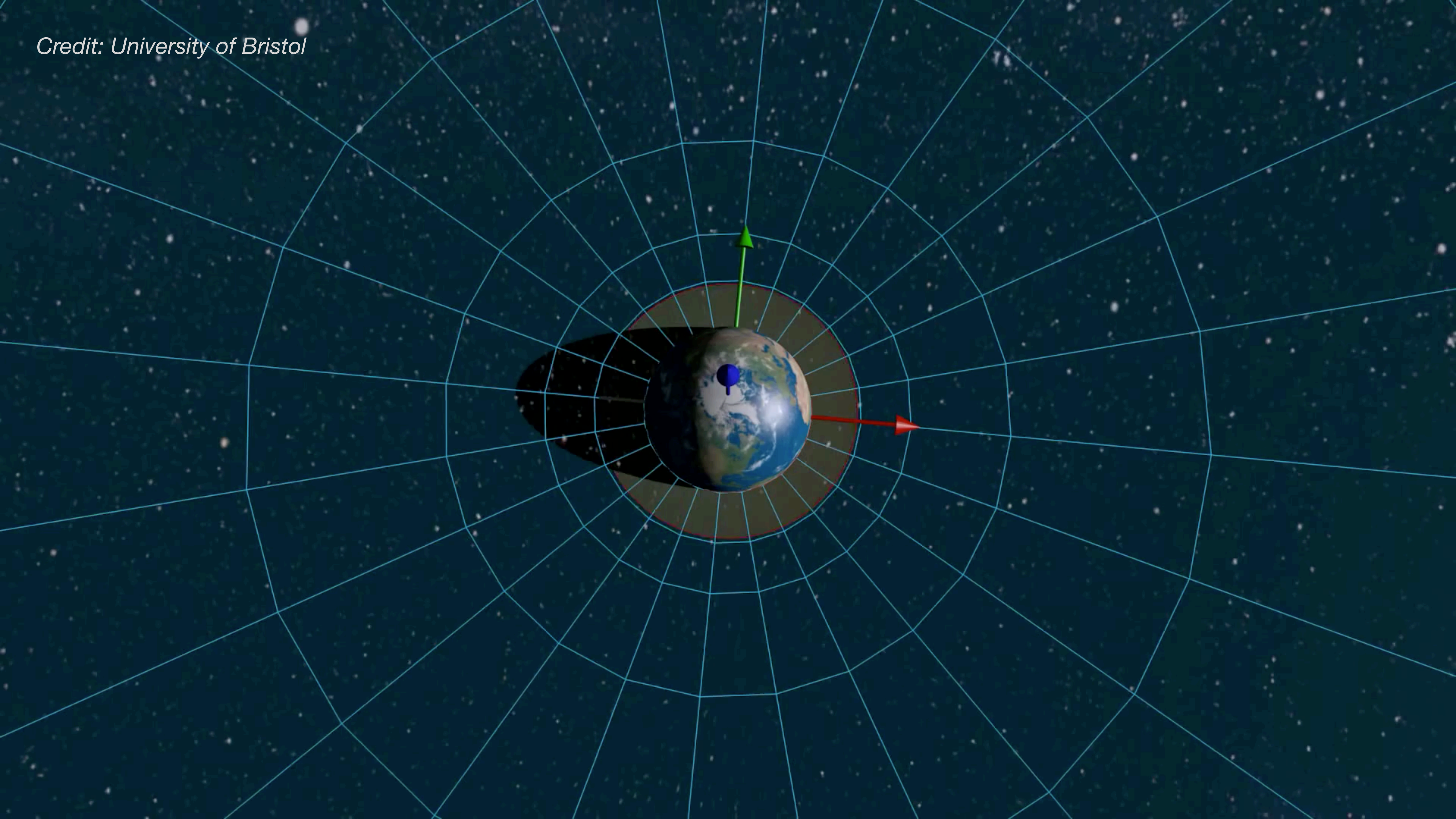
**Bottleneck!**



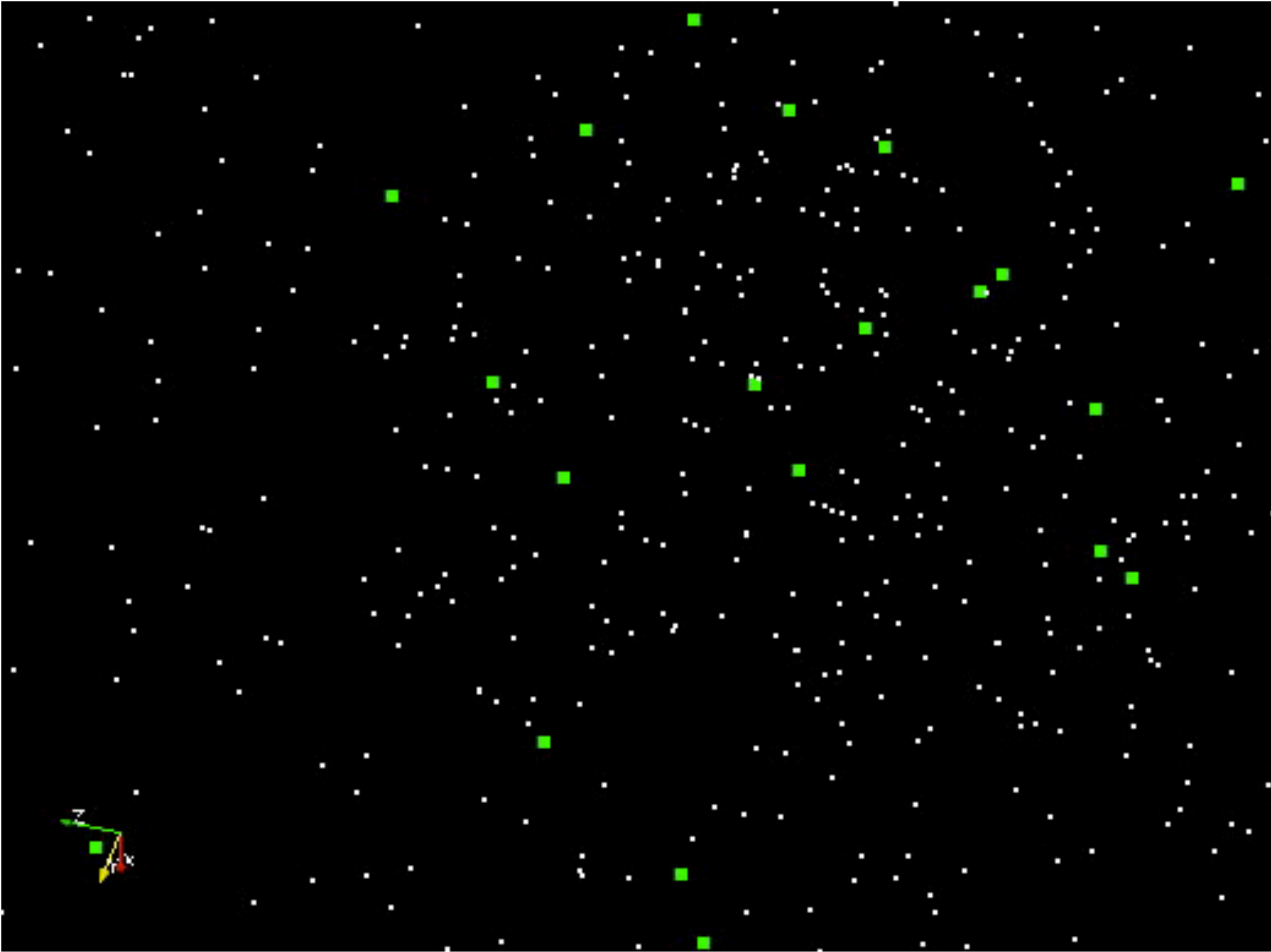
$T < T_{end}$

$T > T_{end}$

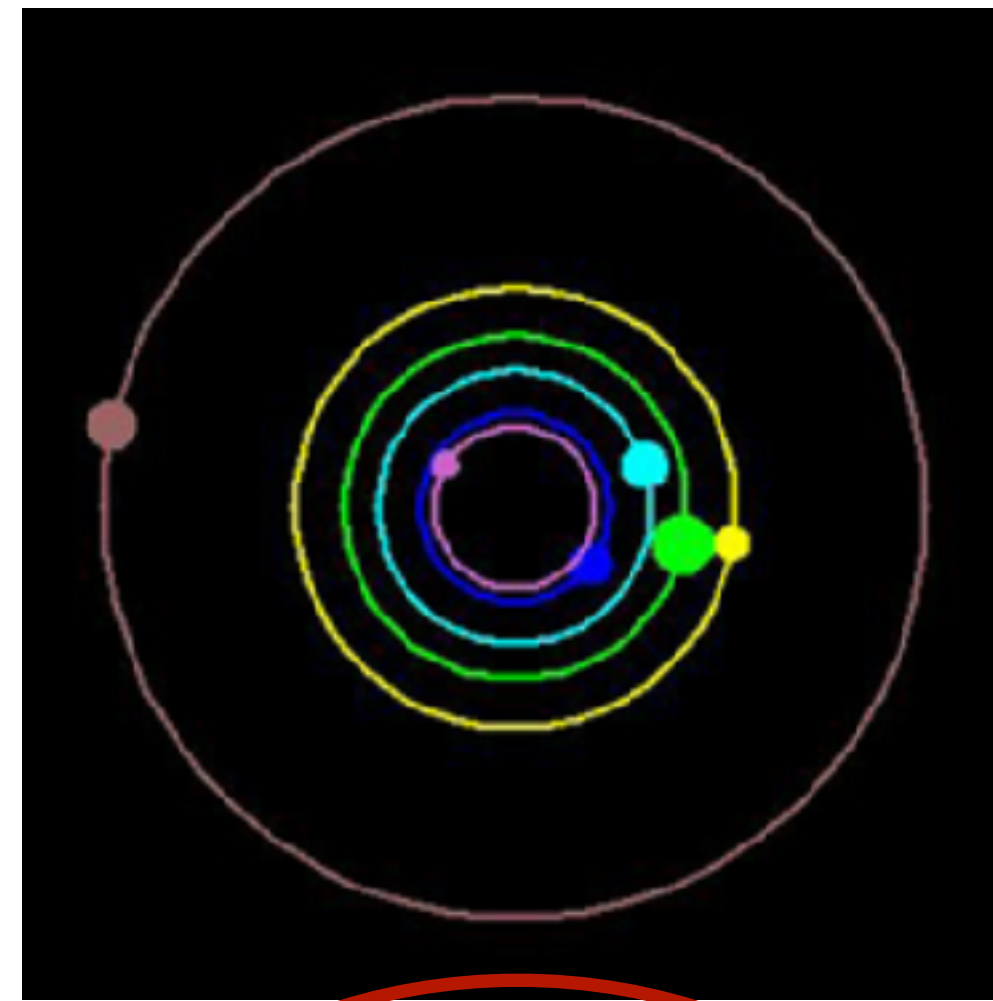
Credit: University of Bristol



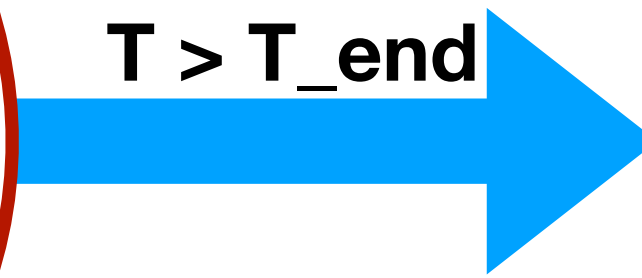
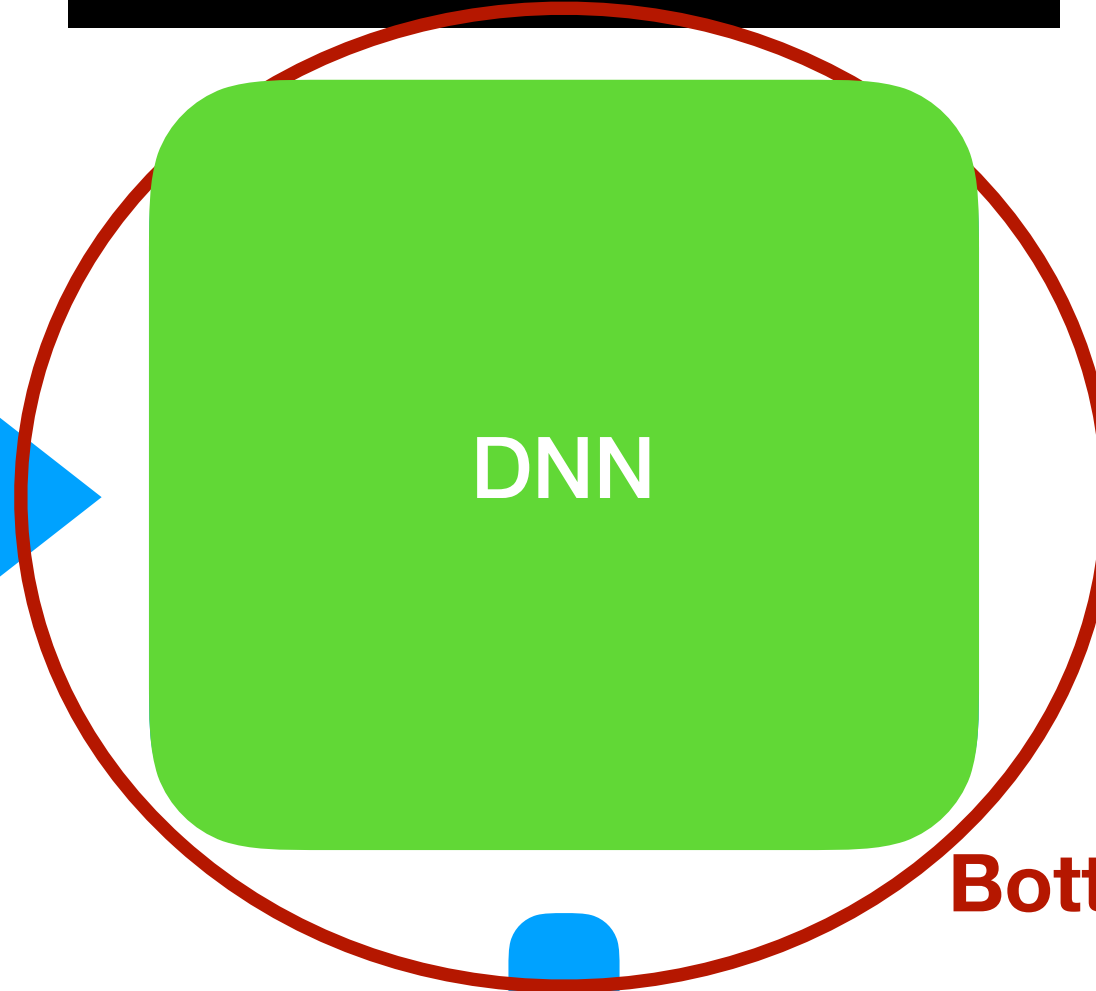
# Coevolution of Planetary Systems and the Host Cluster



# Multi-scale Modeling



Evolve Star clusters  
(dt ~ 1000 yr)

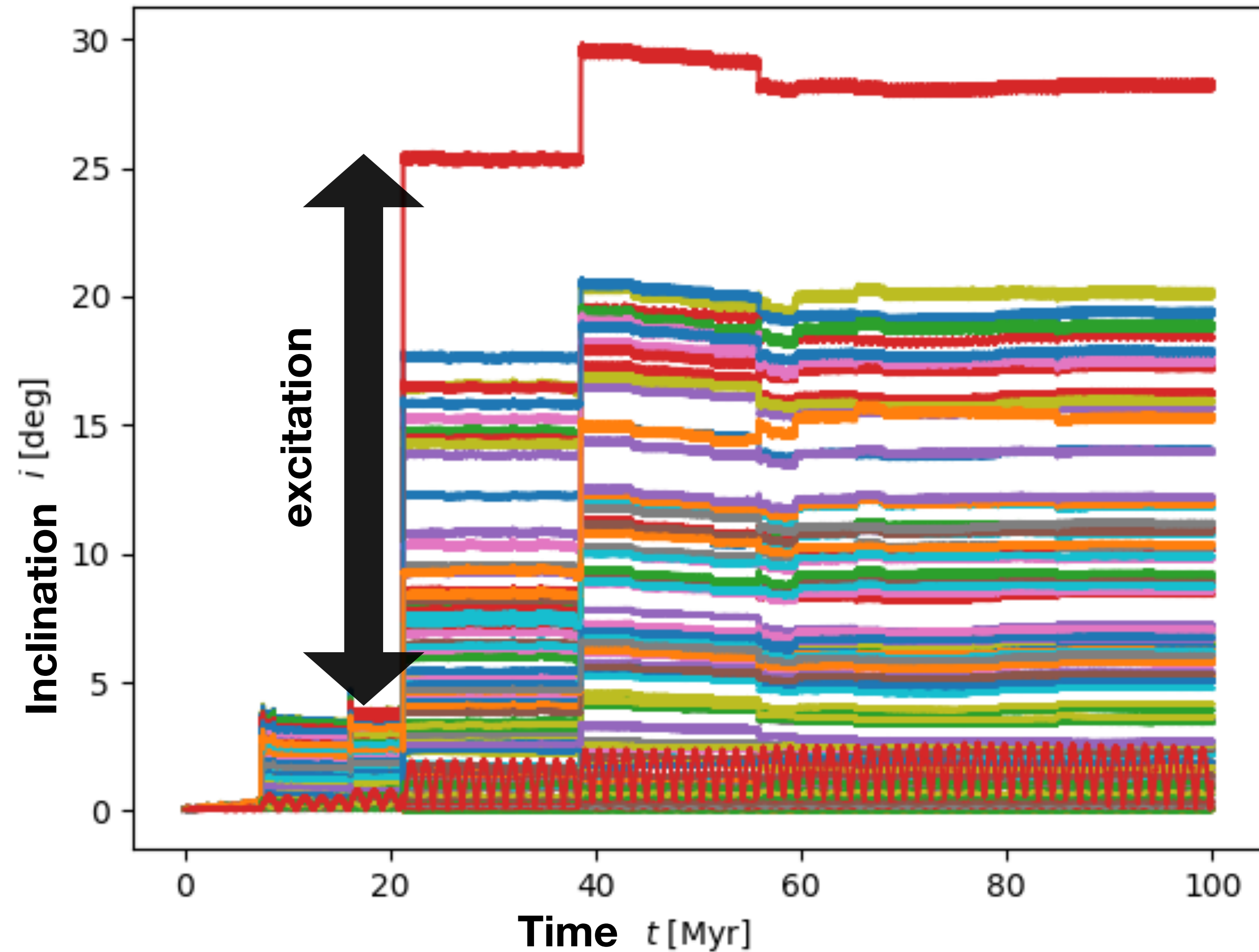


Done



# Can a DNN learn to predict orbits?

*Cai et al. in prep.*



## Challenges

- Predict on extremely long timescales
- The systems exhibit chaotic behaviors
- High dynamic range
- Huge parameter space
- Imbalance training samples — interesting events are rare

# Can a DNN learn to predict orbits?

*Cai et al. in prep.*

Predict individual systems accurately

Very challenging on long timescales

Predict overall statistics accurately

Possible, but simple ML might be enough

Predict both individual systems and overall statistics accurately

Very challenging on long timescales

## Challenges

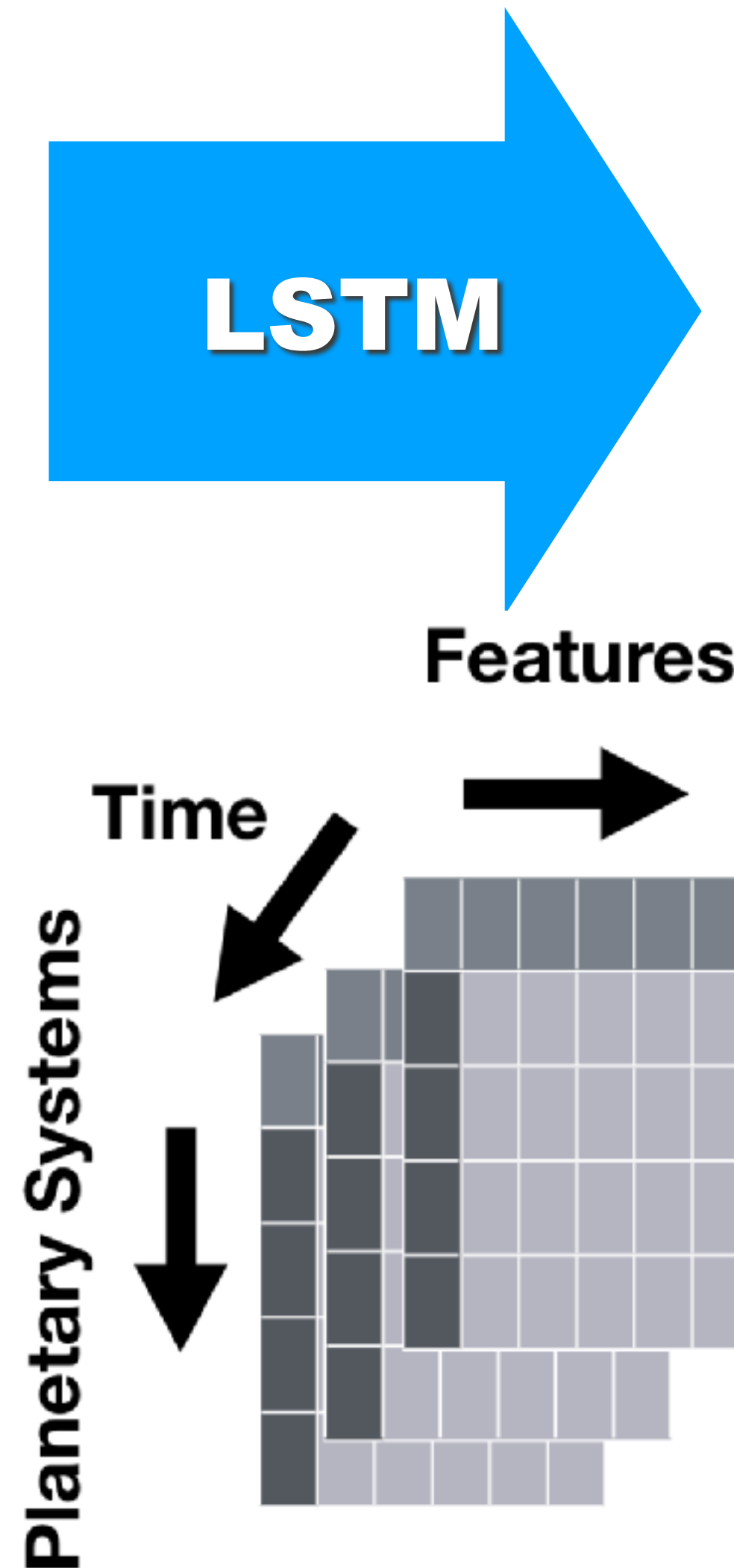
- Predict on extremely long timescales
- The systems exhibit chaotic behaviors
- High dynamic range
- Huge parameter space
- Imbalance training samples — interesting events are rare

# Multivariate Time Series Prediction

*LSTM: Long Short-term Memory (Hochreiter & Schmidhuber 1997)*

## Multiple Features

- Eccentricities
- Inclinations
- Semi-major axis
- Mass of perturber
- distance of the perturber
- velocity of the perturber
- position of the perturber
- .....

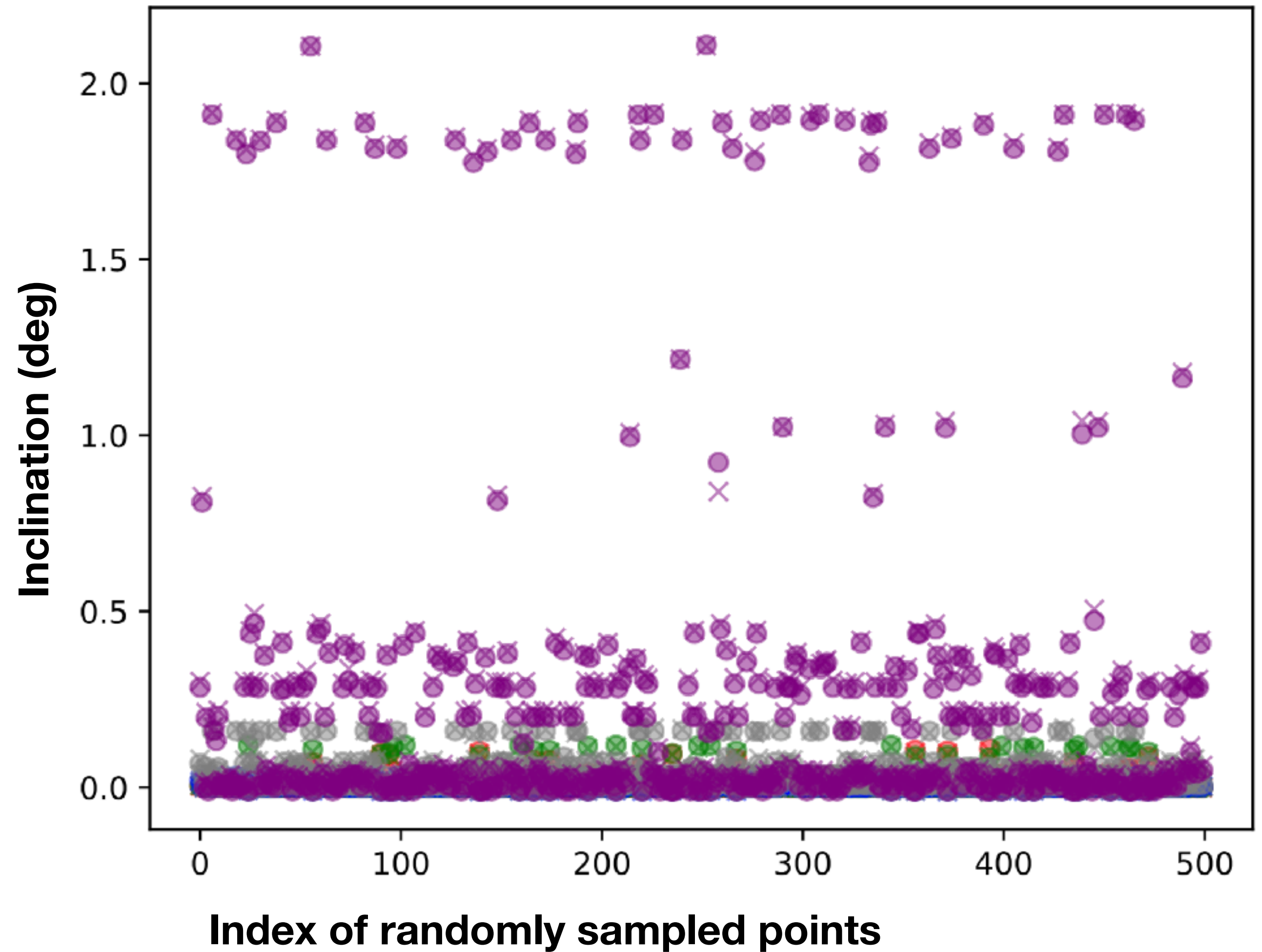


## Next $n$ steps

- Eccentricities
- Inclinations
- Semi-major axis

# Limitation of Time Series Prediction

- Reasonably accurate for short timescales
- Errors accumulate over long timescales

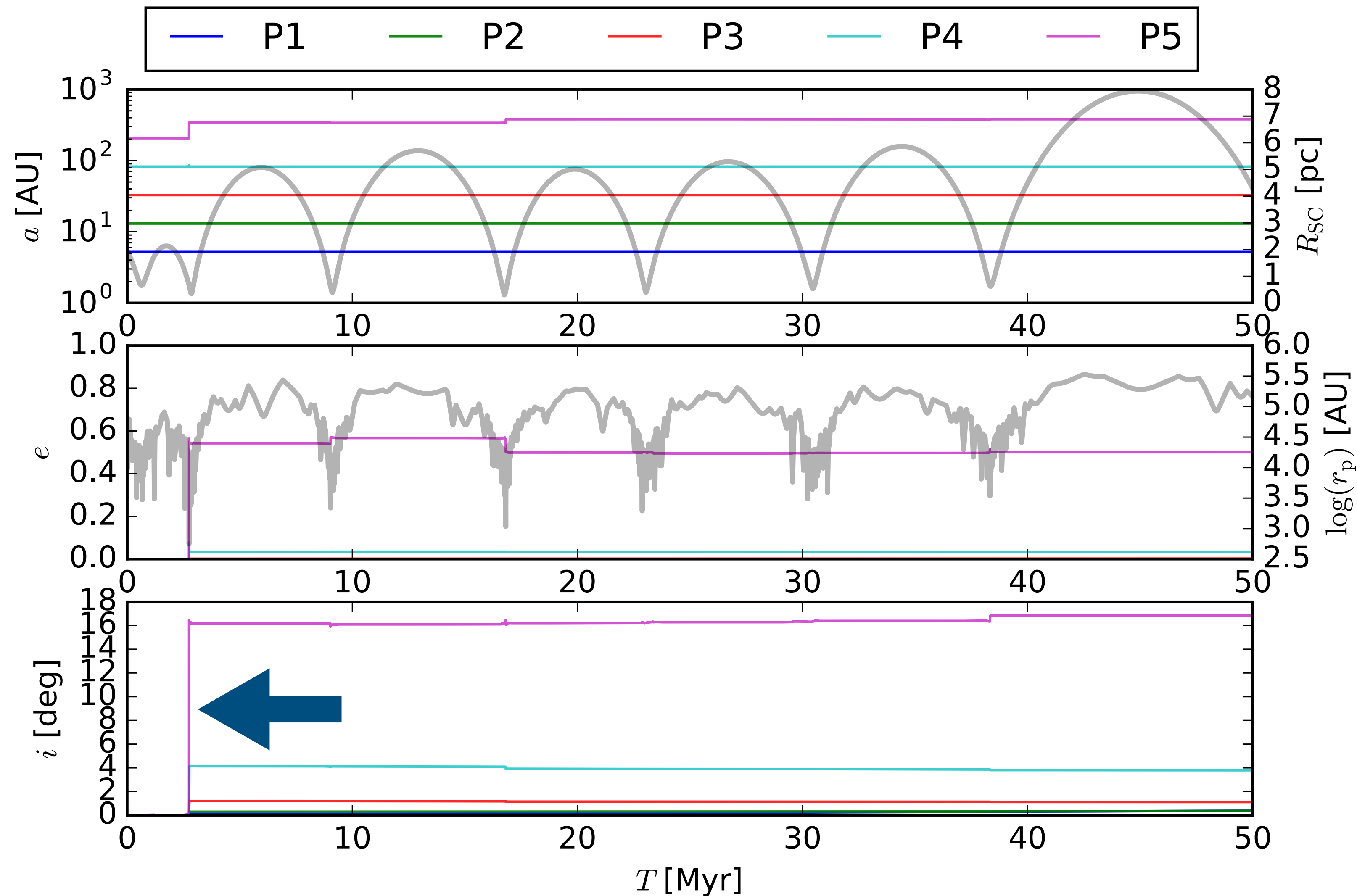




# Supervised learning

## Supervised learning requires:

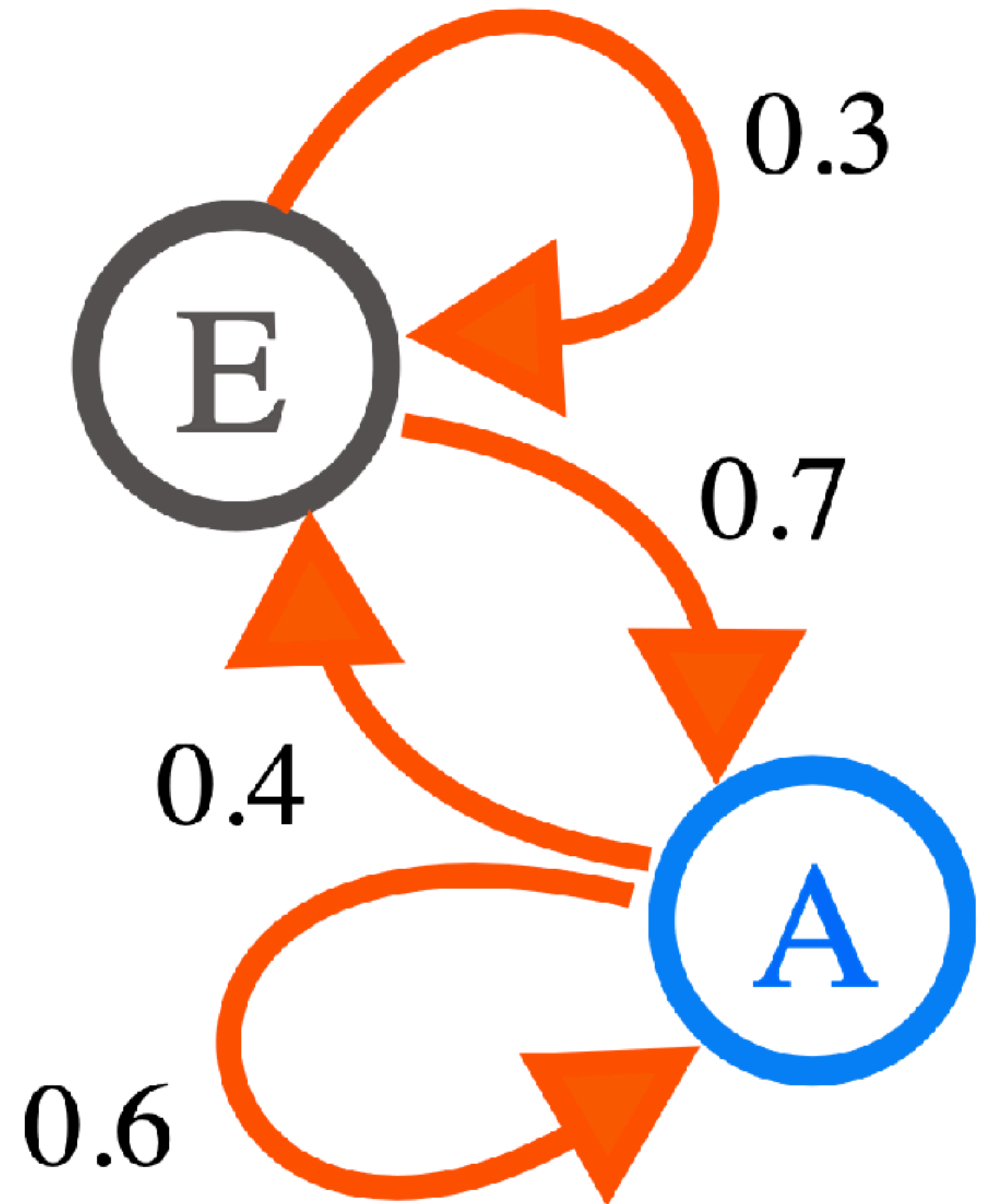
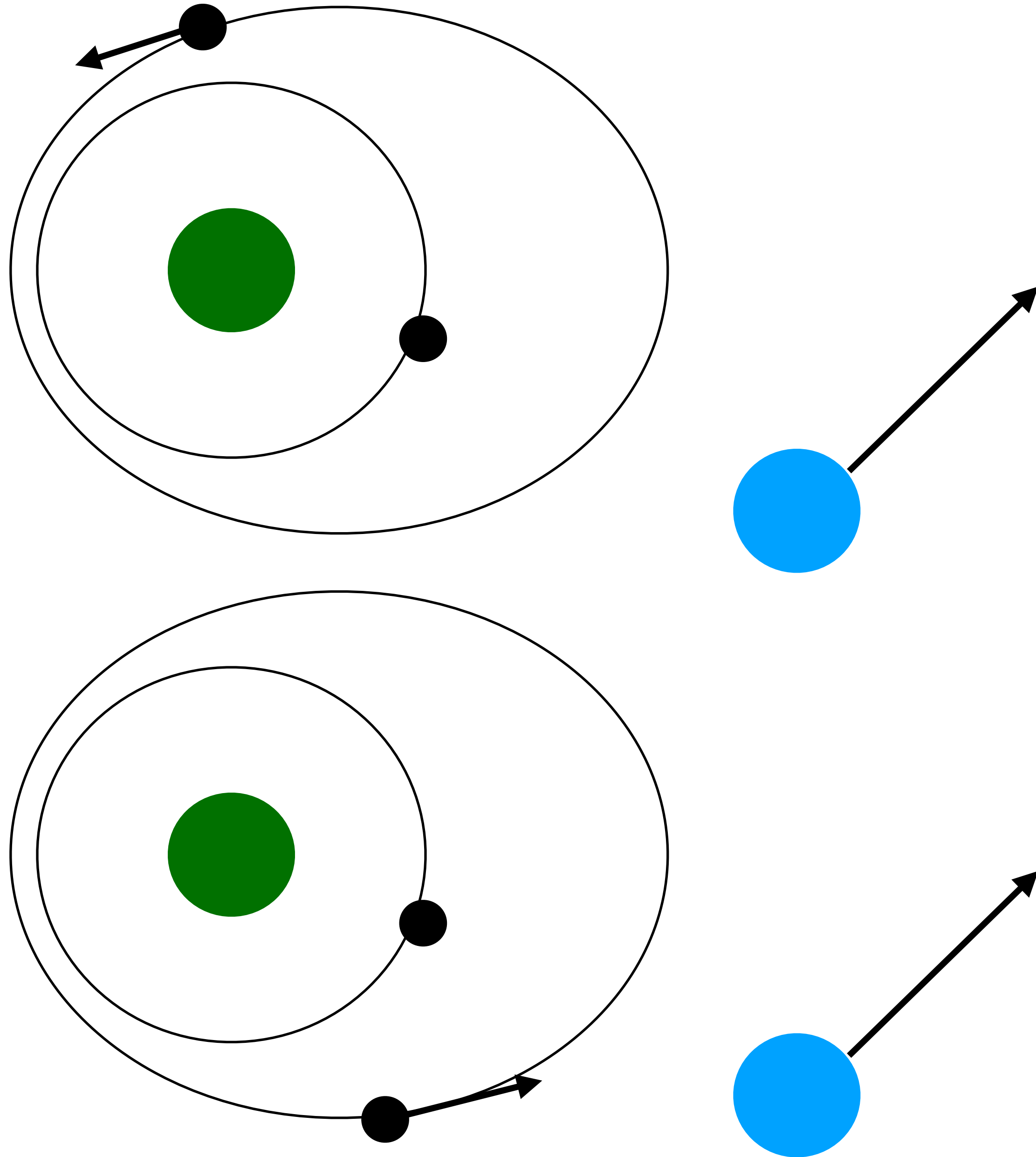
- ➔ Samples are randomized among batches
- ➔ Each batch has the same or similar distribution
- ➔ Samples are independent of each other in the same batch



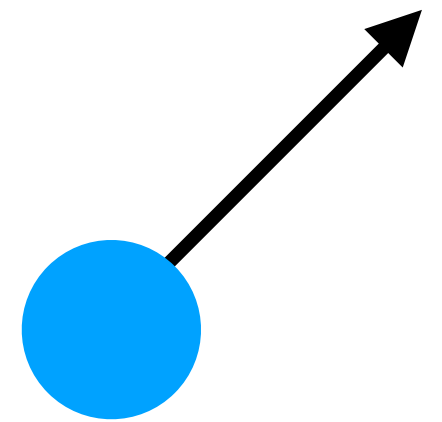
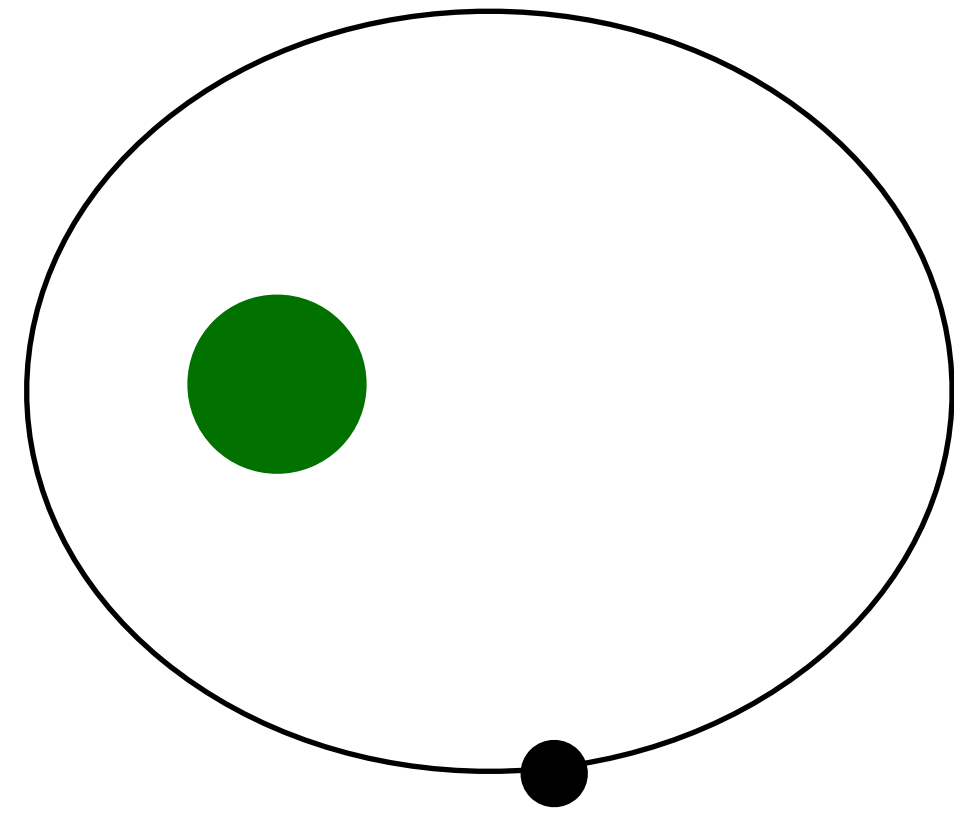
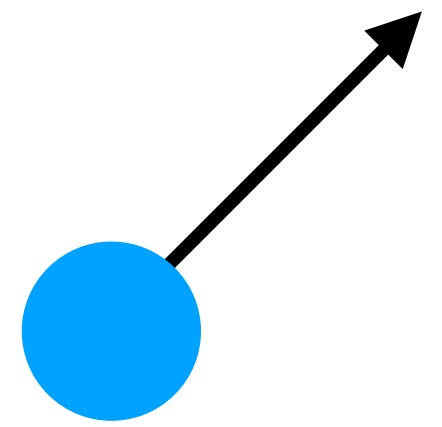
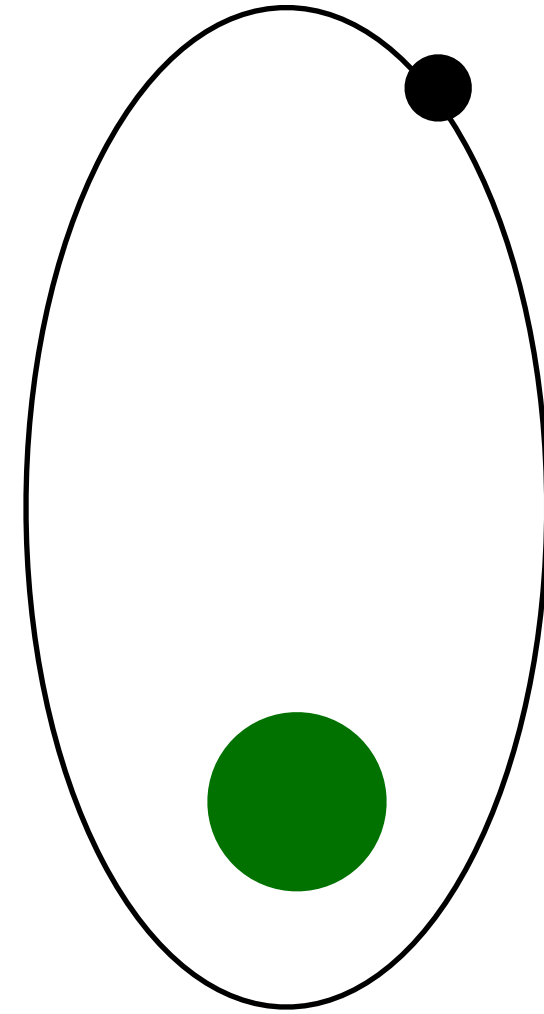
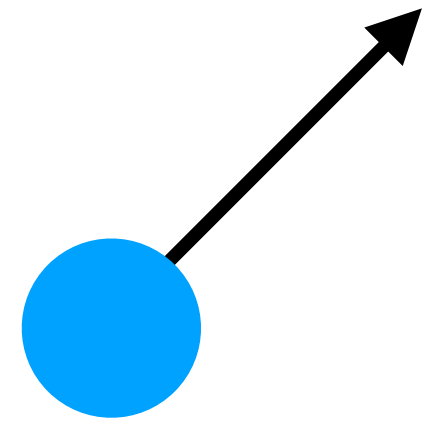
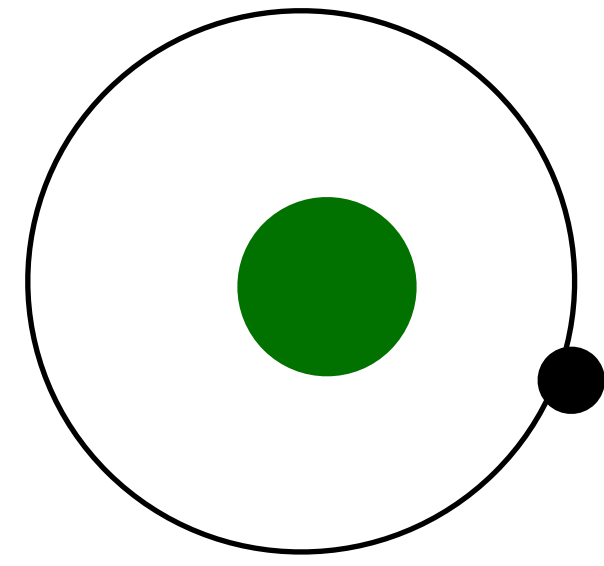
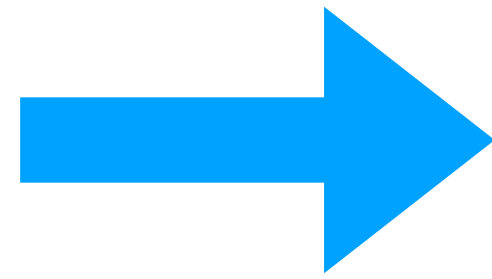
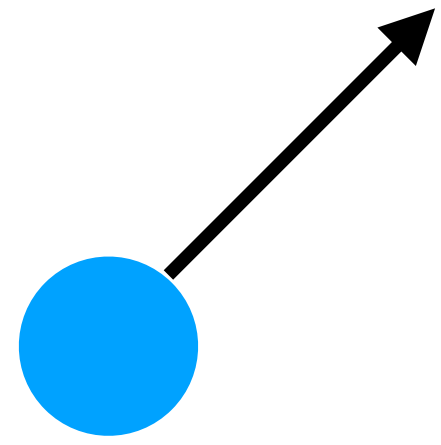
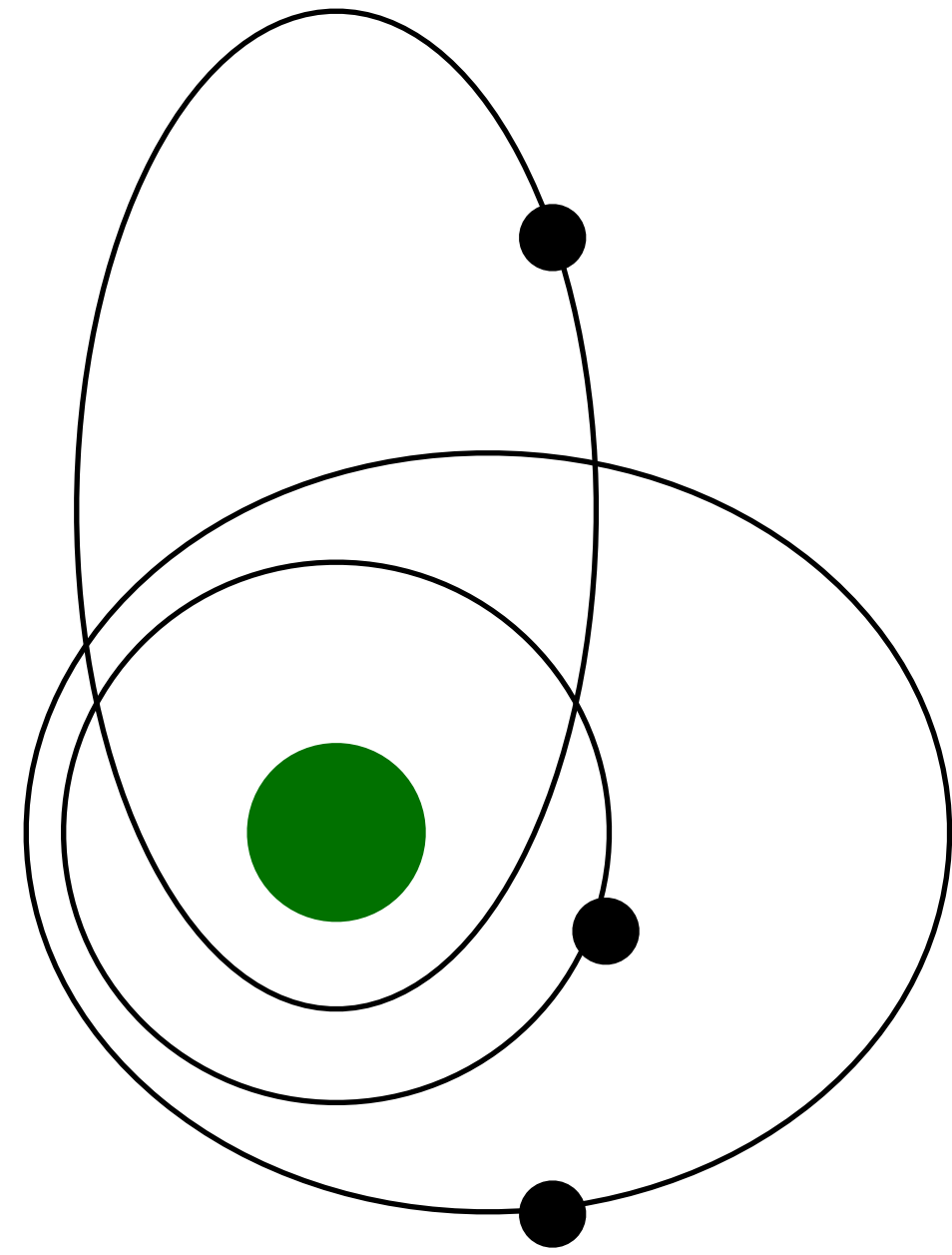
# Better solution?

*Can a neural network learn the physics by itself?*

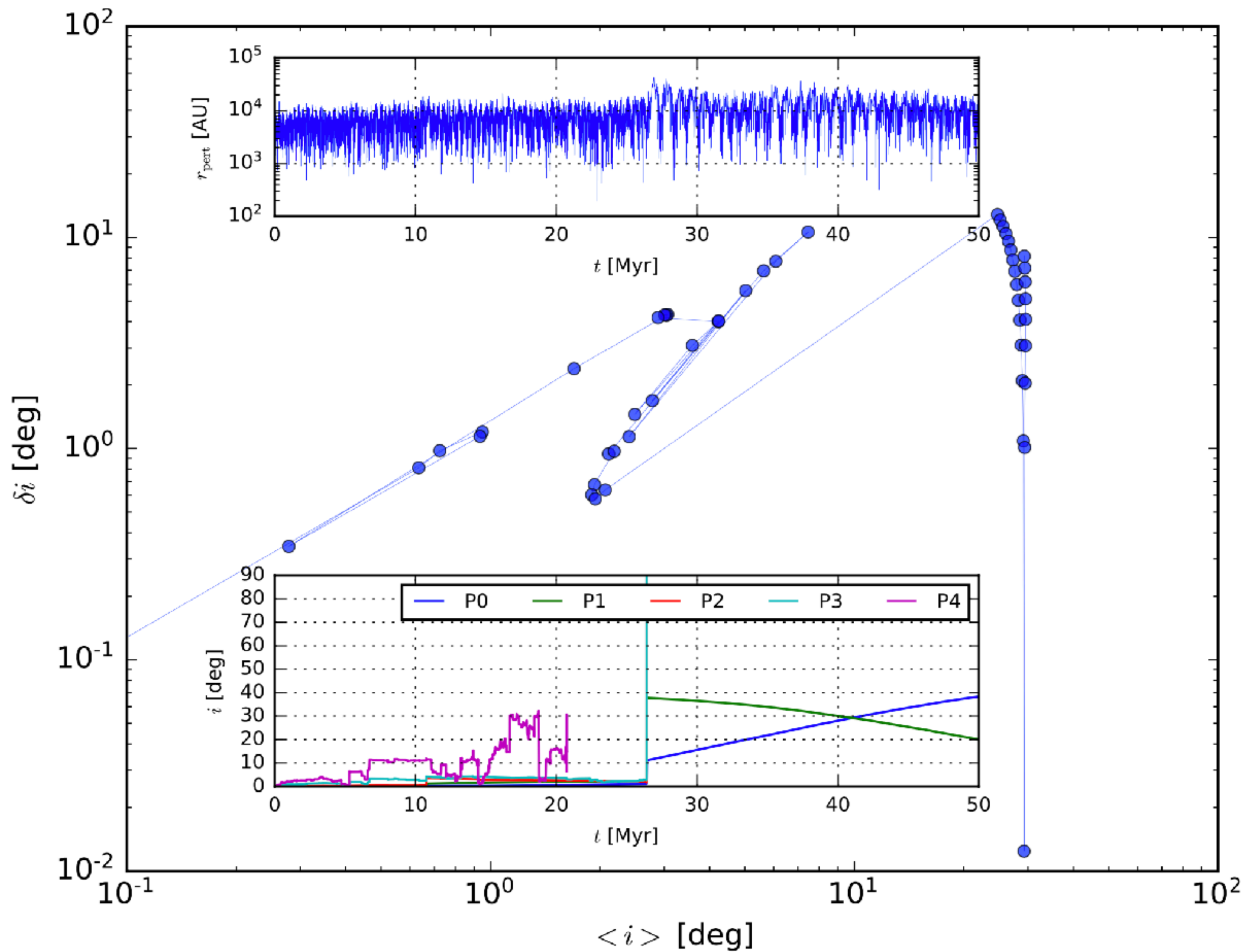
# Stochastic Orbital Changes



Markov chain. Source: Wikipedia



# Can a DNN predict phase-space trajectories?



# Keras-rl library

keras-rl / keras-rl

Used by 206

Watch 217

Star 3,902

Fork 989

Code

Issues 16

Pull requests 23

Projects 1

Wiki

Security

Insights

Deep Reinforcement Learning for Keras. <http://keras-rl.readthedocs.io/>

keras

tensorflow

theano

305 commits

Branch: master

New pull



mirraaj Merge pull request

.github

assets

docs

examples

rl

## Available Agents

Name	Implementation	Observation Space	Action Space
DQN	<code>rl.agents.DQNAgent</code>	discrete or continuous	discrete
DDPG	<code>rl.agents.DDPGAgent</code>	discrete or continuous	continuous
NAF	<code>rl.agents.NAFAgent</code>	discrete or continuous	continuous
CEM	<code>rl.agents.CEMAgent</code>	discrete or continuous	discrete
SARSA	<code>rl.agents.SARSAgent</code>	discrete or continuous	discrete

# OpenAI Gym

openai / gym

Used by 5,592

Watch 939

Star 17,111

Fork 4,610

Code

Issues 148

Pull requests 17

Projects 0

Wiki

Security

Insights



A toolkit for developing and comparing reinforcement learning algorithms. <https://gym.openai.com/>

1,036 commits

78 branches

21 releases

180 contributors

View license

Branch: master

New pull request

Create new file

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 mor-katz and pzhokhov Modify MultiDiscrete comment regarding NOOP (#1537)

Latest commit c03ec69 2 days ago

<a href="#">bin</a>	fix mujoco-related build failure	8 months ago
<a href="#">docs</a>	remove instructions for adding new environments to gym (#1458)	2 months ago
<a href="#">examples</a>	cleanup examples/scripts/sim_env, make it python3 compatible	4 months ago
<a href="#">gym</a>	Modify MultiDiscrete comment regarding NOOP (#1537)	2 days ago
<a href="#">scripts</a>	Cleanup, removal of unmaintained code (#836)	last year
<a href="#">tests/gym/wrappers</a>	Implement FilterObservationWrapper (#1500)	26 days ago

# Modeling with Reinforcement Learning

**Trajectory**  $\tau = (s_0, a_0, s_1, a_1, s_2, a_2, \dots, s_H, a_H, s_{H+1})$

**Reward**  $R(\tau) = r_1 + r_2 + r_3 + \dots + r_H + r_{H+1}$

**Expectation**  $U(\theta) = \sum_{\tau} P(\tau, \theta) R(\tau)$

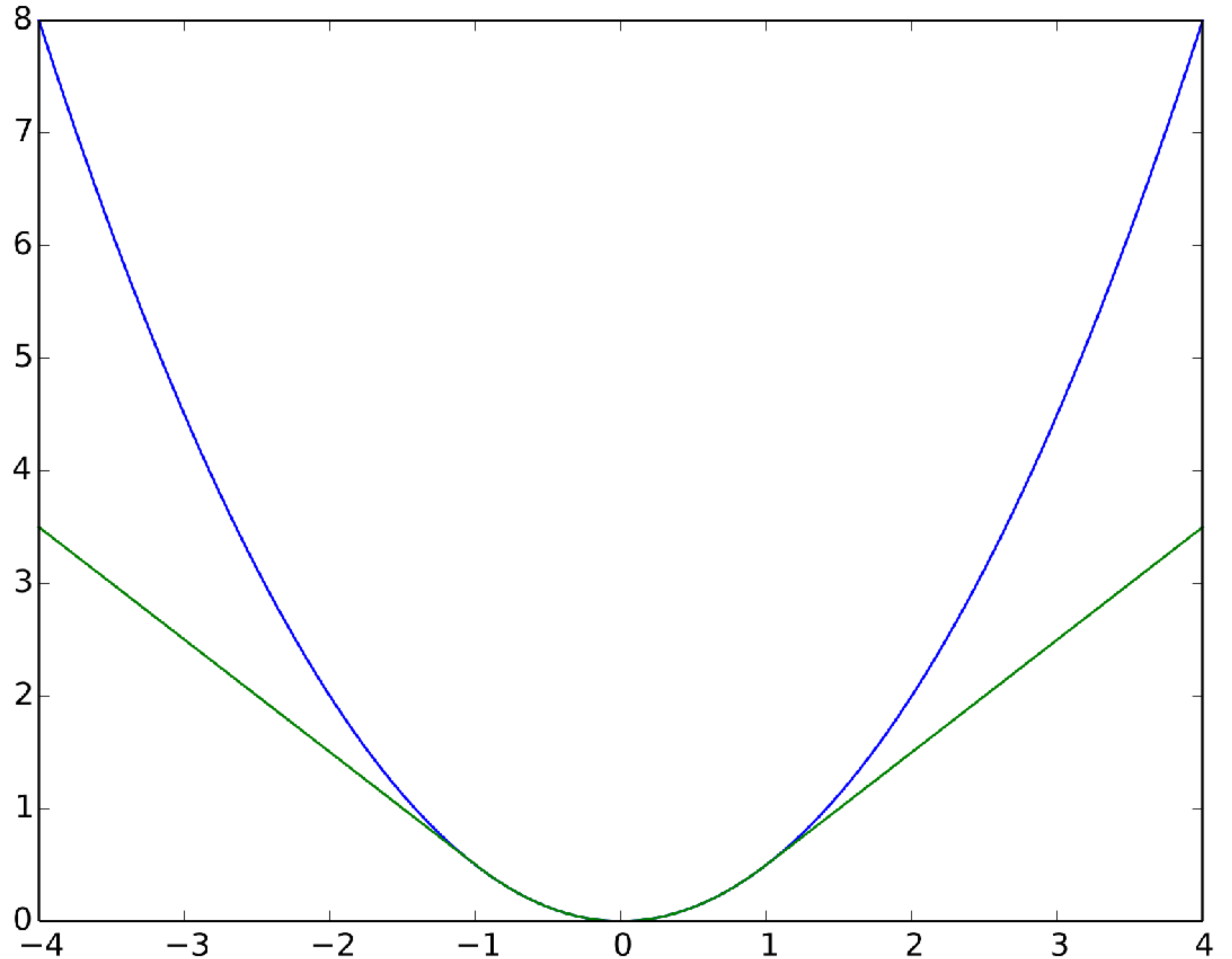
**Gradient**  $\nabla_{\theta} \approx \hat{g} := \frac{1}{m} \sum_{i=1}^m \sum_{t=0}^H \nabla_{\theta} \log \pi_{\theta} (a_t^{(i)} | s_t^{(i)}) R(\tau^{(i)})$

**Update**  $\theta \leftarrow \theta + \alpha \hat{g}$



# Huber Loss

$$L_{\delta}(a) = \begin{cases} \frac{1}{2}a^2 & \text{for } |a| \leq \delta, \\ \delta(|a| - \frac{1}{2}\delta), & \text{otherwise.} \end{cases}$$



# CONTINUOUS CONTROL WITH DEEP REINFORCEMENT LEARNING

**Timothy P. Lillicrap\*, Jonathan J. Hunt\*, Alexander Pritzel, Nicolas Heess,  
Tom Erez, Yuval Tassa, David Silver & Daan Wierstra**

Google Deepmind  
London, UK

{countzero, jjhunt, apritzel, heess,  
etom, tassa, davidsilver, wierstra} @ google.com

## Algorithm 1 DDPG algorithm

Randomly initialize critic network  $Q(s, a|\theta^Q)$  and actor  $\mu(s|\theta^\mu)$  with weights  $\theta^Q$  and  $\theta^\mu$ .  
Initialize target network  $Q'$  and  $\mu'$  with weights  $\theta^{Q'} \leftarrow \theta^Q, \theta^{\mu'} \leftarrow \theta^\mu$   
Initialize replay buffer  $R$   
**for** episode = 1,  $M$  **do**  
    Initialize a random process  $\mathcal{N}$  for action exploration  
    Receive initial observation state  $s_1$   
    **for**  $t = 1, T$  **do**  
        Select action  $a_t = \mu(s_t|\theta^\mu) + \mathcal{N}_t$  according to the current policy and exploration noise  
        Execute action  $a_t$  and observe reward  $r_t$  and observe new state  $s_{t+1}$   
        Store transition  $(s_t, a_t, r_t, s_{t+1})$  in  $R$   
        Sample a random minibatch of  $N$  transitions  $(s_i, a_i, r_i, s_{i+1})$  from  $R$   
        Set  $y_i = r_i + \gamma Q'(s_{i+1}, \mu'(s_{i+1}|\theta^{\mu'})|\theta^{Q'})$   
        Update critic by minimizing the loss:  $L = \frac{1}{N} \sum_i (y_i - Q(s_i, a_i|\theta^Q))^2$   
        Update the actor policy using the sampled policy gradient:

$$\nabla_{\theta^\mu} J \approx \frac{1}{N} \sum_i \nabla_a Q(s, a|\theta^Q)|_{s=s_i, a=\mu(s_i)} \nabla_{\theta^\mu} \mu(s|\theta^\mu)|_{s_i}$$

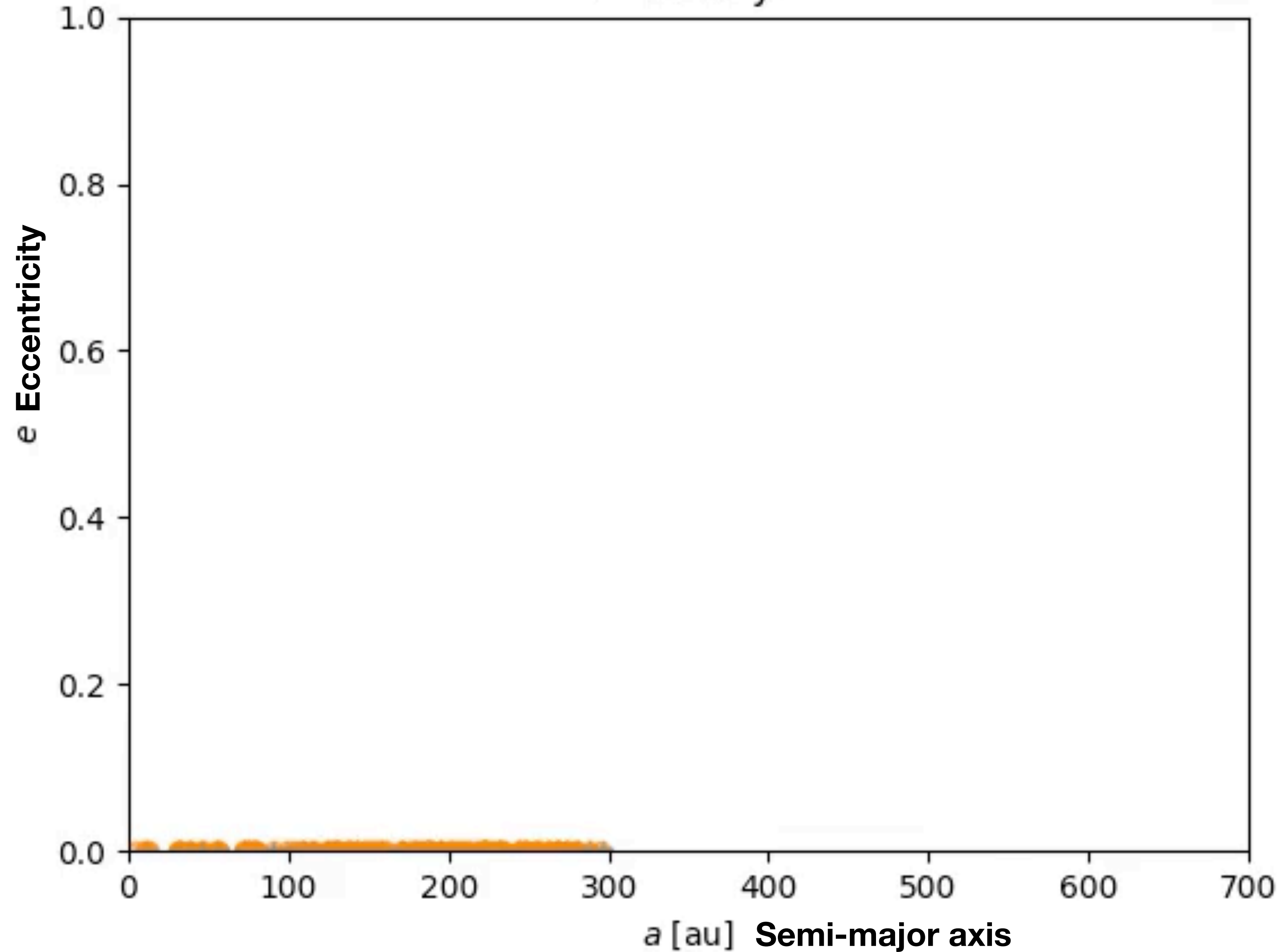
Update the target networks:

$$\begin{aligned}\theta^{Q'} &\leftarrow \tau \theta^Q + (1 - \tau) \theta^{Q'} \\ \theta^{\mu'} &\leftarrow \tau \theta^\mu + (1 - \tau) \theta^{\mu'}\end{aligned}$$

**end for**  
**end for**

# Comparing with N-body simulations

$t = 1000 \text{ yr}$



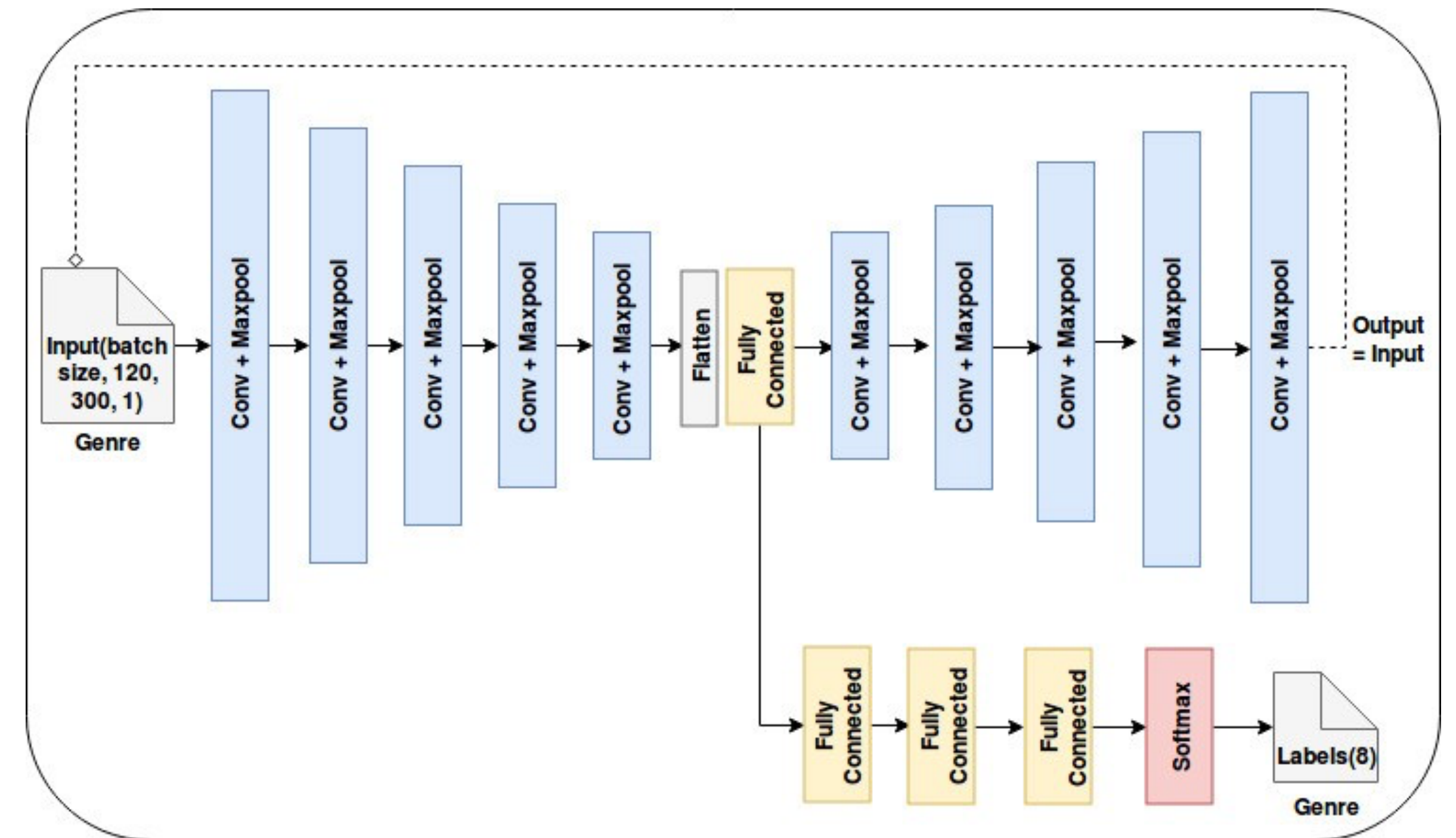
**DQN managed to capture the basic physics**  
**Speed up by a factor of  $10^2 - 10^6$**

**Predict accuracy depends on the resolution**  
**of the training data**

# Predict the future according to the past: pattern recognition

1	4	7								
2	5	...								<b>Semi-major axes</b>
3	6									
1	4	7								
2	5	...								<b>Eccentricities</b>
3	6									
1	4	7								
2	5	...								<b>Inclinations</b>
3	6									
1	4	7								
2	5	...								<b>Perturber distances</b>
3	6									

Variational Autoencoder as classifier

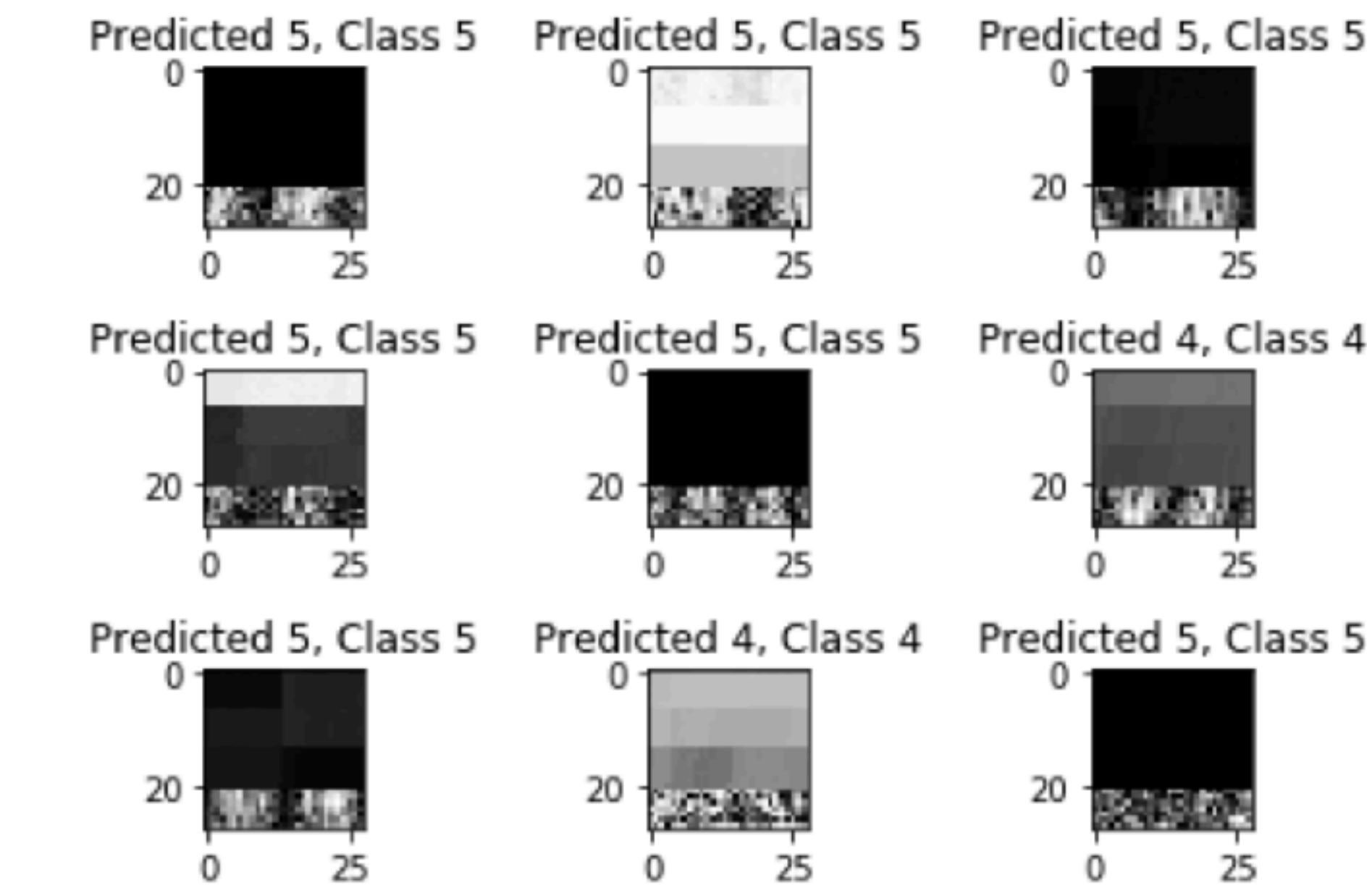
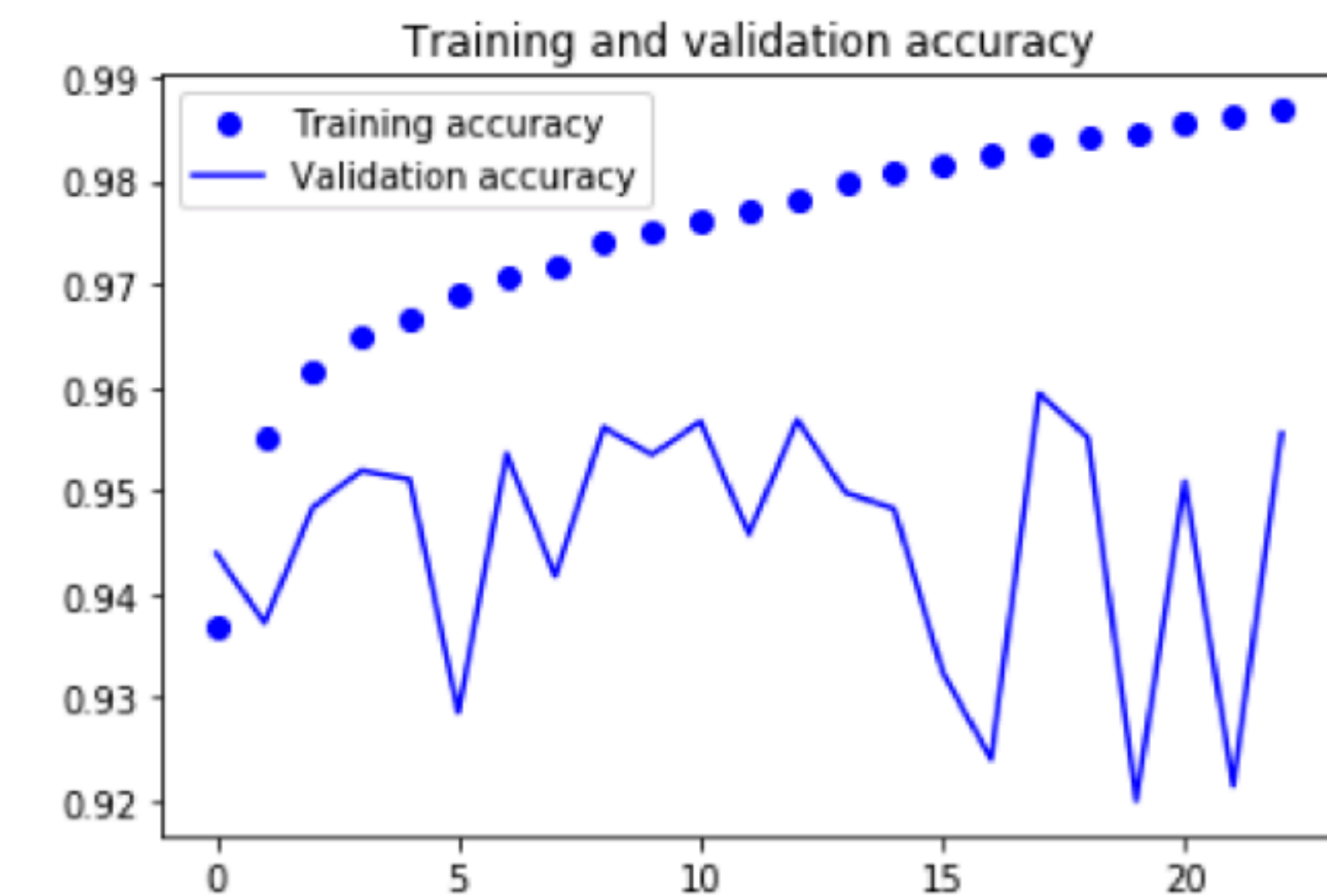
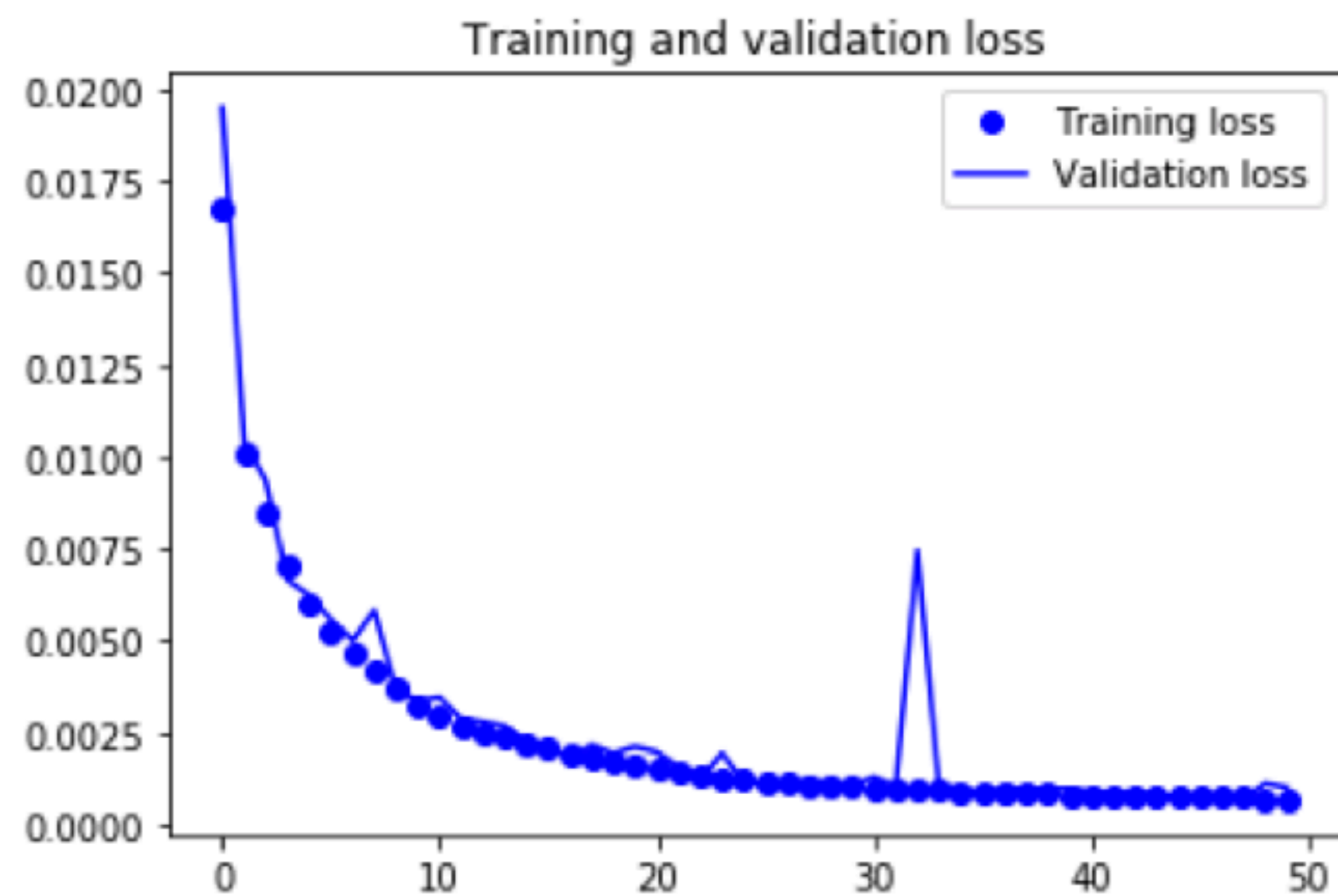


min

max

**Label: the change of orbital eccentricity in the next 1 Myr**

# Predict the future according to the past: pattern recognition



	precision	recall	f1-score	support
Class 0	0.97	0.98	0.98	918
Class 1	0.74	0.76	0.75	141
Class 2	0.63	0.83	0.72	252
Class 3	0.76	0.71	0.73	897
Class 4	0.93	0.92	0.93	24992
Class 5	0.97	0.98	0.97	71386
Class 6	0.83	0.80	0.81	1471
Class 7	0.74	0.80	0.77	440
Class 8	0.85	0.74	0.79	221
Class 9	0.89	0.89	0.89	162
micro avg	0.96	0.96	0.96	100880
macro avg	0.83	0.84	0.83	100880
weighted avg	0.96	0.96	0.96	100880

# Summary



## Challenges:

- Underlying systems chaotic
- High dynamic range
- Extremely imbalance training samples
- Extremely long term prediction needed
- System not deterministic

## Conclusions

- Supervised learning is useful, but only for short-term prediction
- It is unusual to use RL for time series prediction, but it seems that RL can indeed learn physical laws
- Long term error inevitable, because we can't change the chaotic nature of the systems
- Multiple neural network architectures needed to collectively tackle the problem
- DL/RL can be useful for multi-scale modeling in physics

**Bonus slides**



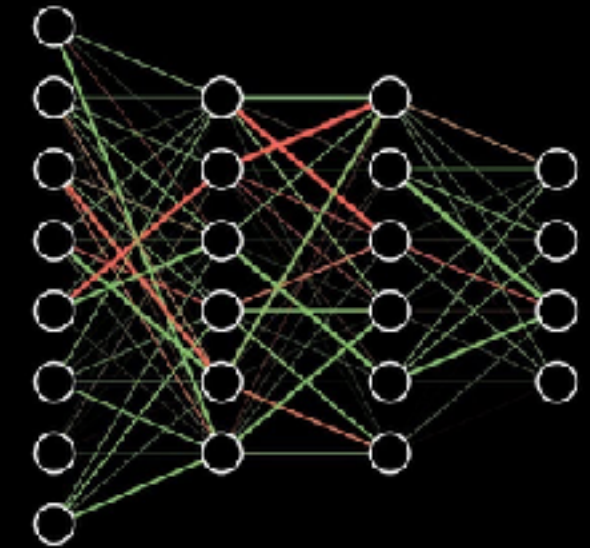
# Galaxy merger Simulations

**Simulation**

**Observation**

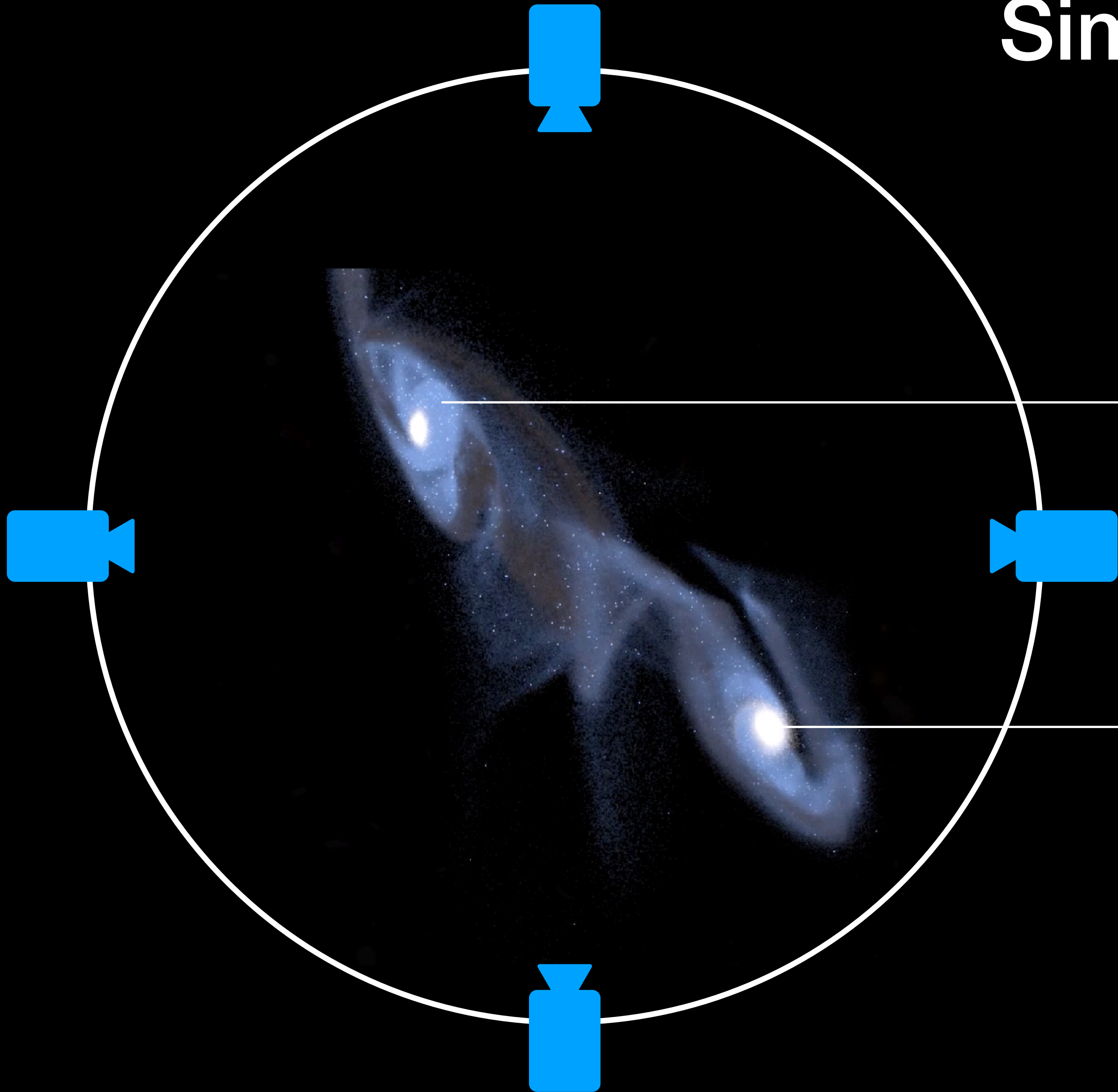


Neural Networks



**Astrophysical problem → Pattern recognition problem**

# Simulations & Visualization



$M_1$

= 1:1, 1:2, 1:3, 2:3, ...

$M_2$

~60 GB of image data

A lot more simulation data